

# Celestial optical theorem and its applications

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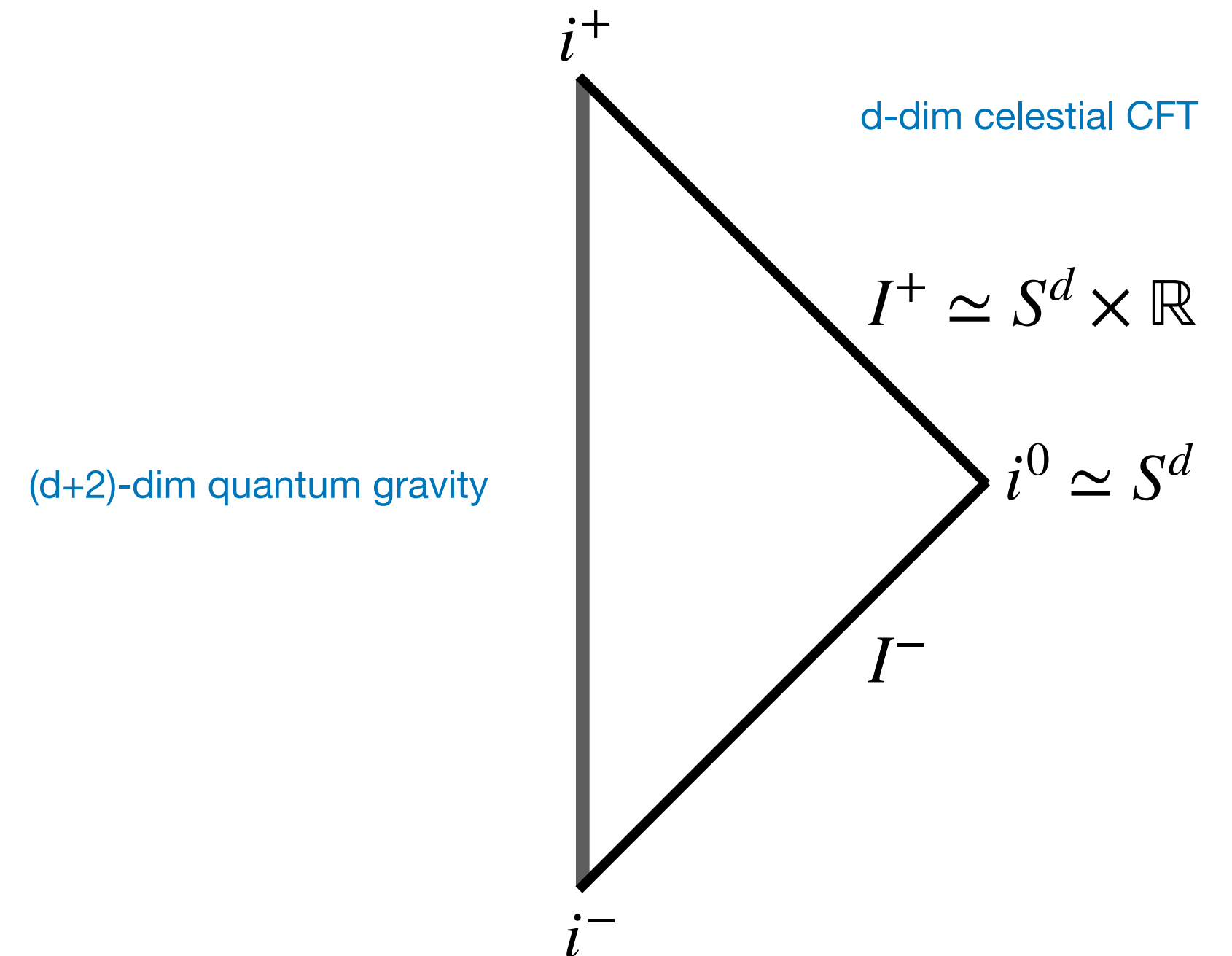
2024 06

- Chi-Ming Chang, RL, and Wen-Jie Ma. Split representation in celestial holography. 11 2023. arXiv:2311.08736.
- RL and Wen-Jie Ma. Massive celestial amplitudes and celestial amplitudes beyond four points. 4 2024. arXiv:2404.01920.
- RL and Wen-Jie Ma. Celestial Optical Theorem. 4 2024. arXiv:2404.18898.

# Introduction

# Celestial holography

- Proposal:
  - bulk:  $(d+2)$ -dim quantum gravity on AF spacetime;
  - bdry: putative  $d$ -dim CCFT on the celestial sphere.
- Symmetry:
  - bulk: Lorentz symmetry  $SO(d+1,1)$ ;
  - bdry: Euclidean conformal symmetry  $SO(d+1,1)$ ;
  - here we do not discuss asymptotic symmetries/large gauge transformations.
- Observables:
  - bulk: scattering amplitudes  $\mathcal{M}$ ;
  - bdry: celestial amplitudes  $\mathcal{A}$  as conformal correlators.
- We focus on  $d = 2$ , bosonic objects:  $(\Delta, J) = (h + \bar{h}, h - \bar{h})$  and  $J \in \mathbb{Z}$ .
- We try to learn **universal properties** of CCFT from examples: **S-matrix unitarity**.



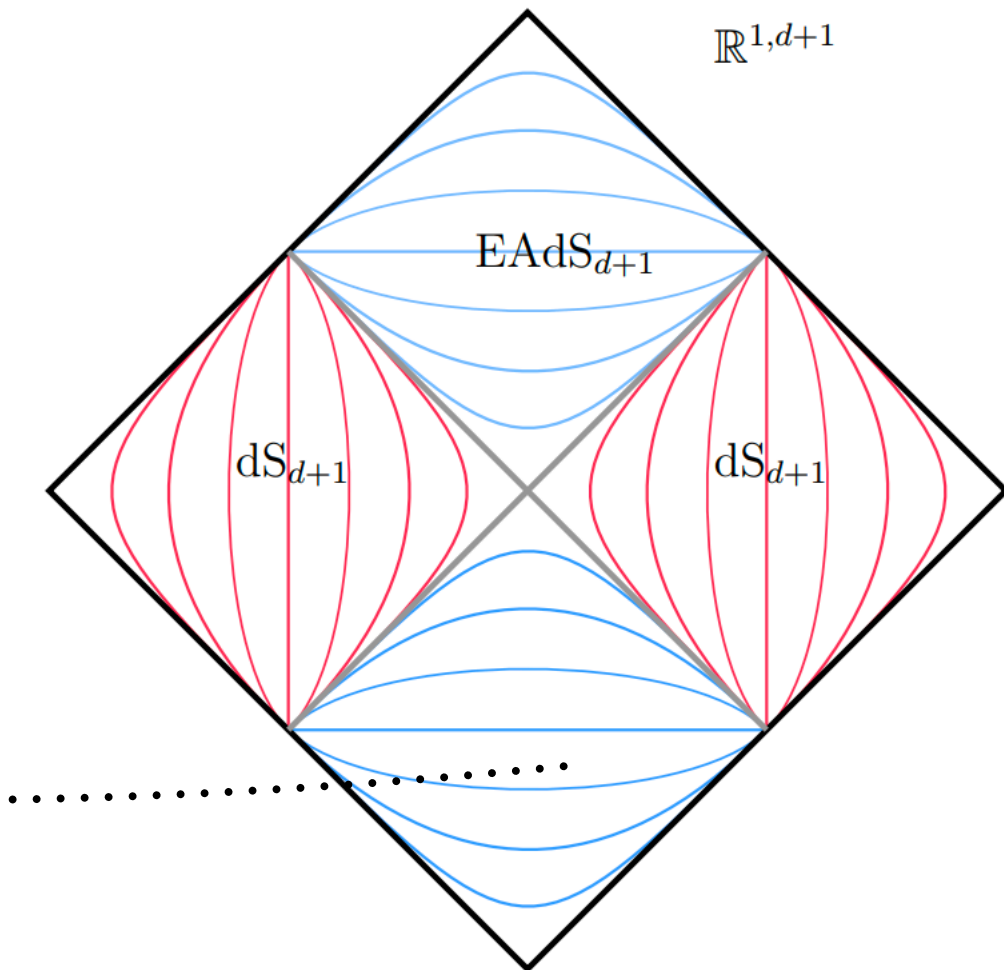
# Dictionary

[Pasterski, Shao, Strominger]

$$\mathcal{A}_{I \rightarrow F}^{h_I | h_F}(z_I | z_F) = \left( \prod_{i=1}^{n+m} \int \frac{d^3 k_i}{k_i^0} \phi_{h_i, m_i}^{\{\mu\} \ell_i}(z_i, k_i) \right) \mathcal{M}_{I \rightarrow F, \{\mu\} \ell_I \{\mu\} \ell_F}(k_I | k_F),$$

mass-shell integral
conformal basis

celestial amplitudes
scattering amplitudes



Hyperbolic slicing of momentum space

- Continuous complex conformal dimension  $\Delta$ :
  - $\Delta \in 1 + i\mathbb{R}$ ;
  - required by invertibility and completeness.
- Mass and conformal dimension are **not** related.
- Massless particle:
  - the integral transform is reduced to the Mellin transform;
  - collinear singularity of  $\mathcal{M} \Leftrightarrow$  OPE singularity of  $\mathcal{A}$ .

# Conformal basis $\phi(z, k)$

- Defining property:
  - intertwine  $SO(3,1)$ 's linear action on  $\mathbb{R}^{3,1}$  and conformal action on  $S^2$ ;
  - solutions of EOM;
- Can be constructed from AdS bulk-to-bdry propagator;
- Quantum numbers:
  - mass  $m$
  - conformal dimension  $\Delta$
  - conformal spin  $J$
- bulk spin  $\ell$  wrt  $SO(3,1) \rightarrow$  bdry spin  $J$  wrt  $SO(2)$ 
  - massive:  $J \in L = \{-\ell, -\ell + 1, \dots, \ell - 1, \ell\}$
  - massless:  $J \in L = \{-\ell, \ell\}$

Particles (Wigner)

conformal basis

Putative operators (PS)

effectively as primary operators

SOC and inner product are modified. [Crawley, Miller, Narayanan, Strominger'21]

Primary operators (Verma)

# Example: 3-pt scalar celestial amplitudes at tree-level

- 00-m

[Lam,Shao'17]

$$C_{1_0^0+2_0^0\rightarrow 3_{m_3}^0}^{\Delta_1,\Delta_2,\Delta_3} = 2^{-\Delta_{12}} m_3^{\Delta_{12}-d-2} \Gamma\left[\begin{matrix} \frac{\Delta_{13,2}}{2}, \frac{\Delta_{23,1}}{2} \\ \Delta_3 \end{matrix}\right].$$

- m0-m: hypergeometric fun.

[RL, Ma'24]

$$C_{1_{m_1}^0+2_0^0\rightarrow 3_{m_3}^0}^{\Delta_1,\Delta_2,\Delta_3} = d_{\Delta_1,\Delta_2,\Delta_3} 2^{1-\Delta_2} m_1^{-\Delta_{23}} m_3^{-d+\Delta_3} (m_3^2 - m_1^2)^{\Delta_2-1} {}_2F_1\left(\begin{matrix} \frac{\Delta_{23,1}}{2}, \frac{\Delta_{123}-d}{2} \\ \Delta_2 \end{matrix}; \frac{m_1^2 - m_3^2}{m_1^2}\right)$$

- mm-m: triple MB integral or double sum of Wilson poly.

[RL, Ma'24]

$$C_{1_{m_1}^0+2_{m_2}^0\rightarrow 3_{m_3}^0}^{\Delta_1,\Delta_2,\Delta_3} = d_{\Delta_1,\Delta_2,\Delta_3} \frac{(2\pi)^{\frac{d+1}{2}} (m_1 m_2)^{-\frac{d+1}{2}}}{(m_1 + m_2) \Gamma[\frac{d+1}{2}]} \epsilon^{\frac{d-1}{2}} + O(\epsilon^{\frac{d+1}{2}}).$$

- Proportional to the AdS three-point coefficient  $d_{\Delta_1,\Delta_2,\Delta_3}$ .

$$d_{\Delta_1,\Delta_2,\Delta_3} = \frac{1}{2} \pi^{d/2} \Gamma\left[\begin{matrix} \frac{\Delta_{12,3}}{2}, \frac{\Delta_{13,2}}{2}, \frac{\Delta_{23,1}}{2}, \frac{\Delta_{123}-d}{2} \\ \Delta_1, \Delta_2, \Delta_3 \end{matrix}\right]$$

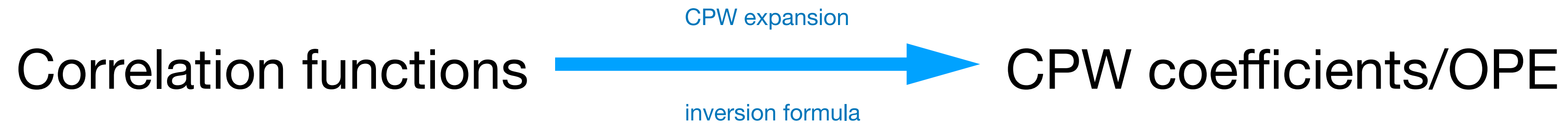
- In AdS/CFT, with  $d_{\Delta_1,\Delta_2,\Delta_3}$  and split representation, double-trace operators appear in OPE.

- What about CCFT?

**celestial amplitudes and OPE**



# Conformal partial wave expansion



$$\mathcal{A}^{h_i}(z_i) = \left( \prod_{j=1}^{n-3} \sum_{J'_j=-\infty}^{+\infty} \int_1^{1+i\infty} \frac{d\Delta'_j}{2\pi i \mu_{h'_j}} \right) \rho_{h'_i}^{h_i} \Psi_{h'_i}^{h_i}(z_i) ,$$

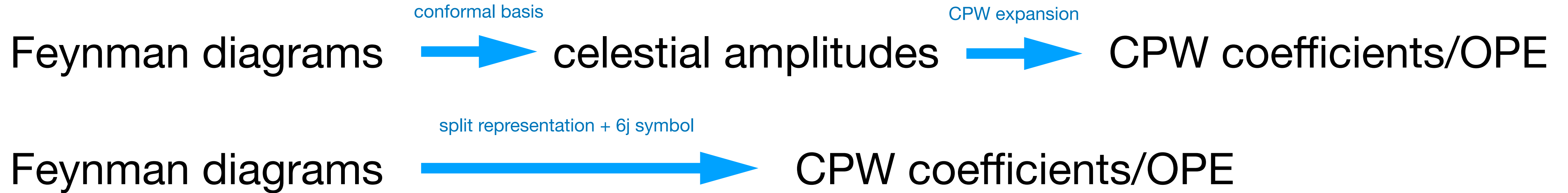
$$\rho_{h'_i}^{h_i} = \frac{1}{\text{Vol SO}(1,3)} \left( \prod_{j=1}^n \int d^2 z_j \right) \left( \Psi_{h'_i}^{h_i}(z_i) \right)^* \mathcal{A}^{h_i}(z_i) .$$

celestial amplitude   integration over reps   CPW coefficient   CPW

- Based on harmonic analysis of  $SO(3,1)$ ;
- Carrying exactly the same dynamical information as celestial amplitude;
  - pole of CPW coefficient  $\Leftrightarrow$  operator in OPE;
  - useful technique for deriving conformal block expansions and OPE.

# Split representation of scalar Feynman propagator

[Chang, RL, Ma'23]



- An efficient tool for computing perturbative celestial amplitudes;
- Based on harmonic analysis of symmetric spaces.

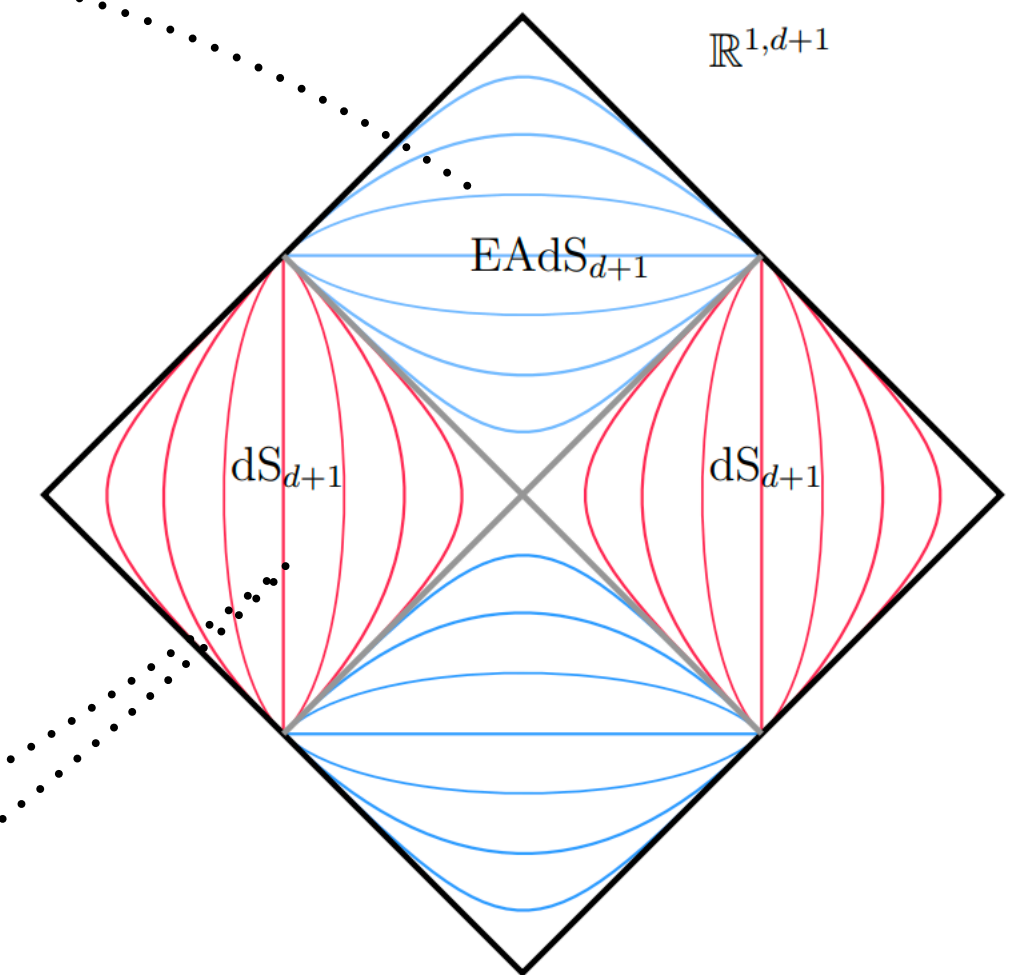
Feynman propagator  $\xrightarrow{\text{integration over mass leaves}}$   $\xrightarrow{\text{integration over reps}}$

$$\begin{aligned} \mathcal{K}_m(X_1, X_2) = & \int_0^{+\infty} \frac{dM}{(2\pi)^{d+2}} \frac{iM^{d+1}}{-M^2 + m^2} \int_{\frac{d}{2}}^{\frac{d}{2}+i\infty} \frac{d\Delta}{2\pi i \mu_\Delta} \int d^d x \left( \phi_{\Delta, M}^- \phi_{\tilde{\Delta}, M}^+ + \phi_{\Delta, M}^+ \phi_{\tilde{\Delta}, M}^- \right) \\ & + \frac{1}{2} \int_0^\infty \frac{dM}{(2\pi)^{d+2}} \frac{iM^{d+1}}{M^2 + m^2} \left( \sum_{\epsilon=0,1} \int_{\frac{d}{2}}^{\frac{d}{2}+i\infty} \frac{(-1)^\epsilon d\Delta}{2\pi i \mu_\Delta} \int d^d x \phi_{\Delta, iM, \epsilon} \phi_{\tilde{\Delta}, iM, \epsilon} \right. \\ & \left. + \alpha_d \sum_{(\Delta, \epsilon) \in D_{\text{dS}}} \text{Res} \frac{(-1)^\epsilon}{\mu_\Delta} \int d^d x \phi_{\Delta, iM, \epsilon} \phi_{d-\Delta, iM, \epsilon} \right). \end{aligned}$$

principal series of  $L^2(\text{EAdS})$

principal series of  $L^2(\text{dS})$

discrete series of  $L^2(\text{dS})$



Hyperbolic slicing of momentum space

# Double-trace operators in celestial OPE

[RL, Ma'24]

- In AdS/CFT, there are double-trace operators in the OPE from tree-level exchange diagrams.

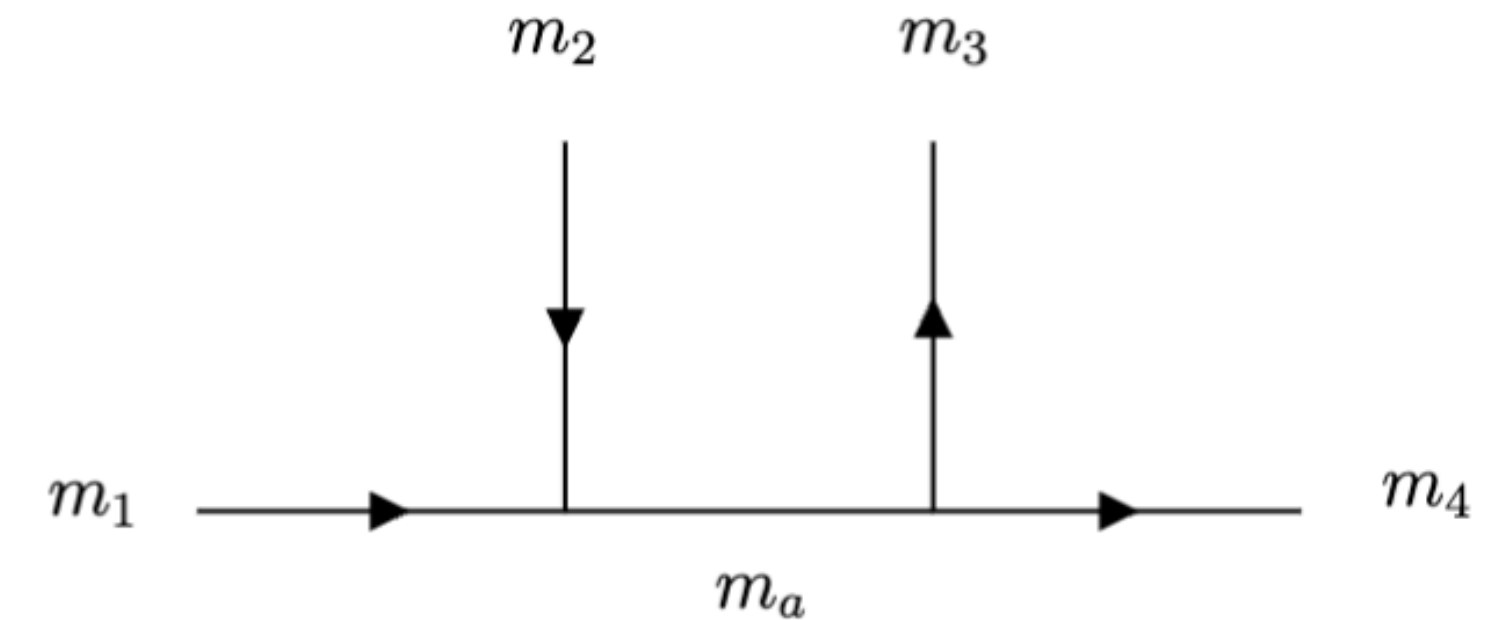
$$\mathcal{O}_{\Delta_1} \mathcal{O}_{\Delta_2} \sim \mathcal{C}_N \mathcal{O}_{\Delta_1 + \Delta_2 + 2N} + \mathcal{C}'_N \mathcal{O}_{\Delta_3 + \Delta_4 + 2N}$$

- In CCFT, we checked that if at most 2-of-5 masses are turned on, there are only double-trace exchanged.

- 00000:

$$\sigma_{\Delta_a}^{\Delta_i} = -i \pi^{-\frac{3d}{2}-2} 2^{-\Delta_{1234}-d-2} 2\pi \delta_c (i(\Delta_{1234} - d - 4))$$

$$\times \Gamma \left[ \begin{matrix} \Delta_3, \Delta_4, \frac{2d-\Delta_{34a}}{2}, \frac{\Delta_{1a,2}}{2}, \frac{\Delta_{2a,1}}{2}, \frac{\Delta_{a,34+d}}{2}, \frac{\Delta_{3a,4}}{2}, \frac{\Delta_{4a,3}}{2} \\ d - \Delta_3, d - \Delta_4, \frac{\Delta_{34,a}}{2}, \Delta_a, \Delta_a - \frac{d}{2}, \frac{\Delta_{34a}-d}{2} \end{matrix} \right].$$



- m0000

$$\sigma_{\Delta_a}^{\Delta_i} = -i \pi^{-d-2} 2^{-\Delta_{234}-d-3} m^{\Delta_{234}-d-4}$$

$$\times \Gamma \left[ \begin{matrix} \Delta_3, \frac{-\Delta_{1234}+d+4}{2}, \frac{\Delta_{1,234+4}}{2}, \Delta_4, \frac{\Delta_{12,a}}{2}, \frac{2d-\Delta_{34a}}{2}, \frac{\Delta_{1a,2}}{2}, \frac{\Delta_{2a,1}}{2}, \frac{\Delta_{12a}-d}{2}, \frac{\Delta_{a,34+d}}{2}, \frac{\Delta_{3a,4}}{2}, \frac{\Delta_{4a,3}}{2} \\ \Delta_1, d - \Delta_3, d - \Delta_4, \frac{-\Delta_{34a}+d+4}{2}, \frac{\Delta_{34,a}}{2}, \Delta_a, \Delta_a - \frac{d}{2}, \frac{\Delta_{a,34+4}}{2}, \frac{\Delta_{34a}-d}{2} \end{matrix} \right].$$

# Multi-particle operators in celestial OPE

[Ball,Hu,Pasterski'23] [Guevara,Hu,Pasterski'24] [RL,Ma'24]

- New two-particle operator with  $\Delta = \Delta_1 + \Delta_2 + 2\Delta_3 - 4$  were observed by OPE limit analysis. [Ball,Hu,Pasterski'23] [Guevara,Hu,Pasterski'24]
- Necessary for OPE associativity.
- Confirmed in 5,6-point tree diagrams, and there should be infinitely many new types of operators, even at tree-level.

[RL,Ma'24]

$$\begin{aligned}
 \mathcal{O}_1^+ \mathcal{O}_2^+ &\sim \overset{4\text{-pt}}{\mathcal{C}_{\Delta_{12}-2+2N}} \mathcal{O}_{\Delta_{12}-2+2N}^+ + \overset{5\text{-pt}}{\mathcal{C}_{\Delta_{12}+2\Delta_3-4+2N}} \mathcal{O}_{\Delta_{12}+2\Delta_3-4+2N}^+ \\
 &\quad + \underset{6\text{-pt}}{\mathcal{C}_{\Delta_{12}+2\Delta_{34}-6+2N}} \mathcal{O}_{\Delta_{12}+2\Delta_{34}-6+2N}^+ .
 \end{aligned}$$

$$\begin{aligned}
 \sigma_{\Delta_a, \Delta_b, \Delta_c} &= -i \pi^{-\frac{7d}{2}-6} 2^{-\Delta_{123456}-3d-8} 2\pi \delta_c (i(\Delta_{123456} - d - 8)) \\
 &\times \Gamma \left[ \Delta_5, \Delta_6, \frac{\Delta_{1a,2}}{2}, \frac{\Delta_{2a,1}}{2}, \frac{\Delta_{3a,b}}{2}, \frac{\Delta_{3b,a}}{2}, \frac{\Delta_{ab,3}}{2}, \frac{\Delta_{4b,c}}{2}, \frac{\Delta_{5c,6}}{2}, \frac{\Delta_{6c,5}}{2}, \frac{\Delta_{4c,b}}{2}, \frac{\Delta_{bc,4}}{2} \right] \\
 &\quad \left[ d - \Delta_5, d - \Delta_6, \Delta_a, \Delta_a - \frac{d}{2}, \Delta_b, \Delta_b - \frac{d}{2}, \frac{\Delta_{56,c}}{2}, \Delta_c, \Delta_c - \frac{d}{2} \right] \\
 &\times \Gamma \left[ \frac{-\Delta_{3456a}+d+6}{2}, \frac{\Delta_{a,3456}+6}{2}, \frac{-\Delta_{456b}+d+4}{2}, \frac{\Delta_{b,456}+4}{2}, \frac{\Delta_{3ab}-d}{2}, \frac{2d-\Delta_{56c}}{2}, \frac{\Delta_{c,56}+d}{2}, \frac{\Delta_{4bc}-d}{2} \right] \\
 &\quad \left[ \frac{-\Delta_{456b}+d+6}{2}, \frac{\Delta_{b,456}+6}{2}, \frac{-\Delta_{56c}+d+4}{2}, \frac{\Delta_{c,56}+4}{2}, \frac{\Delta_{56c}-d}{2} \right] .
 \end{aligned}$$

# Staggered operators in celestial OPE

[Chang,RL,Ma'23] [Chang,RL,Ma,to appear]

- Former examples are s-channel diagram expanded in s-channel OPE; only EAdS part contributes.
- t-channel diagram in t-channel OPE; dS (continuous and discrete) parts contribute.

$$\begin{aligned}
 & t\mathcal{A}_{1_0^0+2_0^0 \rightarrow 3_0^0+4_0^0}(x_i) \\
 &= \frac{im^{\beta-2}}{2^{5+\beta}(2\pi)^4} \csc\left[\frac{\pi\beta}{2}\right] \left[ 2a_0 G_{\frac{1}{2},\frac{1}{2}}^{\Delta_{13},\Delta_{24}} + \sum_{n=0}^{+\infty} \left( b_n G_{\frac{n+2}{2},\frac{n+2}{2}}^{\Delta_{13},\Delta_{24}} + G_{\text{sta},-\frac{n}{2},\frac{n}{2}+1}^{\Delta_{13},\Delta_{24}} + G_{\text{sta},\frac{n}{2}+1,-\frac{n}{2}}^{\Delta_{13},\Delta_{24}} \right) \right],
 \end{aligned}$$

primary operators

staggered operators

- Why they appear? Relation to celestial diamonds?

$$G_{\Delta}^{\Delta_{13},\Delta_{24}}(\chi) = \chi^{\Delta} \sum_{k=0}^{\infty} \frac{(\Delta - \Delta_{13})_k (\Delta + \Delta_{24})_k}{\Gamma[k+1] (2\Delta)_k} \chi^k$$

$$\begin{aligned}
 G_{\text{sta},-\frac{n}{2}}^{\Delta_{13},\Delta_{24}}(\chi) &= \sum_{k=0}^n \frac{(-1)^k (-\frac{n}{2} - \Delta_{13})_k (-\frac{n}{2} + \Delta_{24})_k (n-k)!}{n! \Gamma[k+1]} \chi^{k-\frac{n}{2}} \\
 &+ \sum_{s=0}^n \sum_{k=0}^{\infty} \frac{(-1)^n (-\frac{n}{2} - \Delta_{13})_{k+n+1} (-\frac{n}{2} + \Delta_{24})_{k+n+1}}{n! \Gamma[k+n+2] \Gamma[k+1] (k+1+s)} \chi^{k+\frac{n+2}{2}}.
 \end{aligned}$$

$$L_0|\mathcal{O}^1\rangle = \Delta|\mathcal{O}^1\rangle,$$

$$L_0|\mathcal{O}^2\rangle = (1 - \Delta)|\mathcal{O}^2\rangle,$$

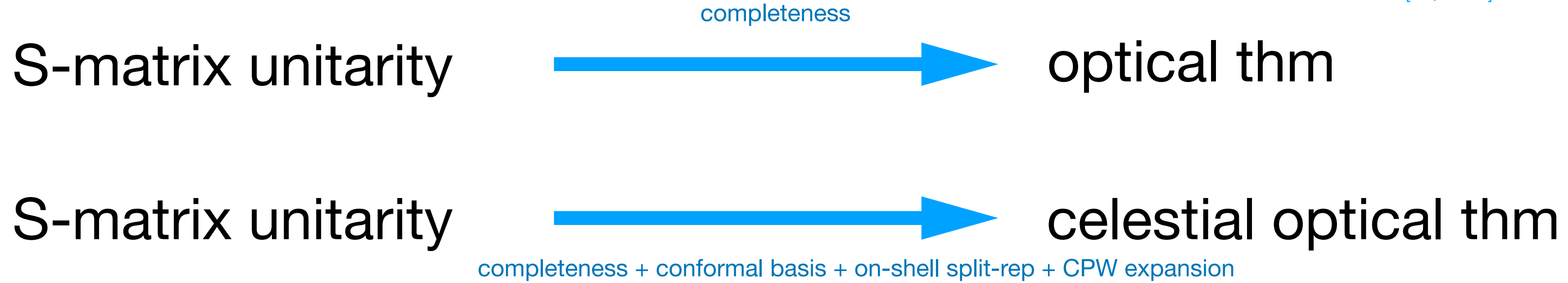
$$L_1|\mathcal{O}^1\rangle = 0,$$

$$L_1|\mathcal{O}^2\rangle = aL_{-1}^{-2\Delta}|\mathcal{O}^1\rangle.$$

# Celestial optical theorem

# Celestial optical theorem

[RL, Ma'24]



$$i\mathcal{M}_{F \rightarrow I}^* - i\mathcal{M}_{I \rightarrow F} = \sum_{n=1}^{+\infty} \sum_{\alpha_n} \int d\Pi_{\alpha_n} \mathcal{M}_{F \rightarrow \alpha_n}^* \mathcal{M}_{I \rightarrow \alpha_n}$$

$$i(\mathcal{A}_{3+4 \rightarrow 1+2}^{\bar{h}_3 \bar{h}_4 | \bar{h}_1 \bar{h}_2})^* - i\mathcal{A}_{1+2 \rightarrow 3+4}^{h_1 h_2 | h_3 h_4} = \sum_{n=1}^{\infty} \sum_{\alpha_n} \frac{1}{S_{\alpha_n}} \left( \prod_{j_{\alpha_n}=1}^n \sum_{J_{j_{\alpha_n}} \in L_{j_{\alpha_n}}} \int_{1-i\infty}^{1+i\infty} \frac{d\Delta_{j_{\alpha_n}}}{2\pi i N_{h_{j_{\alpha_n}}, m_{j_{\alpha_n}}}^{\ell_{j_{\alpha_n}}}} \int d^2 w_{j_{\alpha_n}} \right) \times (\mathcal{A}_{3+4 \rightarrow \alpha_n}^{\bar{h}_3 \bar{h}_4 | h_{\alpha_n}}(z_3, z_4 | w_{\alpha_n}))^* \mathcal{A}_{1+2 \rightarrow \alpha_n}^{h_1 h_2 | h_{\alpha_n}}(z_1, z_2 | w_{\alpha_n})$$

$$S^\dagger S = 1$$

2-2 CPW coefficients

3-pt coefficients

3-pt coefficients in shadow basis

$$i(\rho_{3+4 \rightarrow 1+2, h}^{\bar{h}_3 \bar{h}_4 | \bar{h}_1 \bar{h}_2})^* - i\rho_{1+2 \rightarrow 3+4, h}^{h_1 h_2 | h_3 h_4} = \sum_{\alpha} \frac{\mu_h \sigma_{J, L_\alpha}}{N_{h, m_\alpha}^{\ell_\alpha}} \left( (C_{3+4 \rightarrow \alpha}^{\bar{h}_3 \bar{h}_4 | h})^* C_{1+2 \rightarrow \alpha}^{h_1 h_2 | h} + (C_{3+4 \rightarrow \mathcal{S}[\alpha]}^{\bar{h}_3 \bar{h}_4 | h})^* C_{1+2 \rightarrow \mathcal{S}[\alpha]}^{h_1 h_2 | h} \right) + \sum_{n=2}^{\infty} \sum_{\alpha_n} \frac{1}{S_{\alpha_n}} \left( \prod_{j_{\alpha_n}=1}^n \sum_{J_{j_{\alpha_n}}} \int_{1-i\infty}^{1+i\infty} \frac{d\Delta_{j_{\alpha_n}}}{2\pi i N_{h_{j_{\alpha_n}}, m_{j_{\alpha_n}}}^{\ell_{j_{\alpha_n}}}} \right) \left( \prod_{j=2}^{n-1} \sum_{J'_j=-\infty}^{+\infty} \int_{1-i\infty}^{1+i\infty} \frac{d\Delta'_j}{4\pi i \mu_{h'_j}} \right) (\rho_{3+4 \rightarrow \alpha_n, h'_i}^{\bar{h}_3 \bar{h}_4 | h_{\alpha_n}})^* \rho_{1+2 \rightarrow \alpha_n, h'_i}^{h_1 h_2 | h_{\alpha_n}}$$

internal conformal dimensions

external conformal dimensions

2-n CPW coefficients

# Celestial optical theorem

- Optical thm:
  - unitarity of S-matrix + completeness relation;
  - nonlinear relations between scattering amplitudes;
  - on-shell, nonperturbative.
- Celestial optical thm:
  - unitarity of S-matrix + completeness relation;
  - nonlinear relations between CPW coefficients;
  - on-shell, nonperturbative.
- Optical thm:
  - phase space integrals are hard to manipulate;
  - no natural partial wave expansion for higher-points.
- Celestial optical thm:
  - scalar integral equations of CPW coefficients;
  - natural basis for all points.



# Positivity

- We focus on elastic scattering  $1 + 2 \rightarrow 1 + 2$ .
- Notice:  $\rho_{1+2 \rightarrow 1+2, h}^{h_1 h_2 | h_3 h_4} \equiv \rho_J^{J_i}(\Delta; \Delta_i)$  is a function of  $\Delta$  and  $\Delta_i$  for each fixed spin  $J$  and  $J_i$ .
- Positivity:
  - nonperturbative criterion of CCFT;
  - similar to that in cosmological bootstrap from dS unitarity; [\[Hogervorst, Penedones, Vaziri'21\]](#) [\[Di Pietro, Gorbenko, Komatsu'21\]](#)

$$\text{Im}(\rho_{1+2 \rightarrow 1+2, h}^{h_1 h_2 | \bar{h}_1^* \bar{h}_2^*}) \geq 0, \quad \text{if } \Delta \in 1 + i\mathbb{R}.$$

# Analyticity assumptions

- Analyticity assumptions:
  - $\rho_J^{J_i}(\Delta; \Delta_i)$  is meromorphic with respect to  $\Delta \in \mathbb{C}$  and decays to zero as  $\Delta \rightarrow \infty$ ;
    - natural assumption for CFT;
    - can be loosened to include branch points.
  - $\rho_J^{J_i}(\Delta; \Delta_i)$  is also meromorphic with respect to the external conformal dimensions  $\Delta_i \in \mathbb{C}^4$ ;
    - technical assumption, inspired by perturbative results;
    - can be loosened to smaller domain of  $\Delta_i$ , to include branch points.

# Applications

- Exactly one of the following statements is true:
  - for any  $J$  and  $J_i$ ,  $\rho_J^{J_i}(\Delta; \Delta_i) = 0$ ;
  - for  $J_1 = -J_3$  and  $J_2 = -J_4$ ,  $\rho_J^{J_i}(\Delta; \Delta_i)$  contains at least one  $\Delta$ -pole.
- Given any celestial three-point coefficient  $C_{1+2 \rightarrow \alpha}^{h_1 h_2 | h}$ , if  $\Delta = f(\Delta_1, \Delta_2)$  is a pole of  $C_{1+2 \rightarrow \alpha}^{h_1 h_2 | h}$  for some meromorphic  $f$ , and the intersection between the hypersurface  $\Delta = f(\Delta_1, \Delta_2)$  and the principal series  $\Re \Delta = 1$  is nonempty, then  $\Delta = f(\Delta_1, \Delta_2)$  is a pole of  $\rho_J^{J_i}(\Delta; \Delta_i)$  when  $J_1 = -J_3$  and  $J_2 = -J_4$ .
  - Hence double-trace poles  $\Delta = \Delta_1 + \Delta_2 + \dots$  do not have anomalous dimensions, in contrast to AdS/CFT!
- If the two incoming particles are both massless, exactly one of the statements is true:
  - for any  $J$  and  $J_i$ ,  $\rho_J^{J_i}(\Delta; \Delta_i) = 0$ ;
  - for  $J_1 = -J_3$  and  $J_2 = -J_4$ ,  $\rho_J^{J_i}(\Delta; \Delta_i)$  contains only simple  $\Delta$ -poles in a definite set;
  - in shadow basis, **only** double-trace operators appear in the OPE.
- All the properties are consistent with known perturbative examples.

## **Future directions**

# Future directions

- (Off-shell) Split representation for spinning particles are still incomplete.
- S-matrix crossing symmetry should lead further constraints on CPW coefficients.
- Solving the bootstrap equations from the celestial optical theorem.
- Better understanding of staggered modules (work in progress).

**Thanks for attention!**