# 5D Weyl Double Copy

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- 2 Classical double copy
- 3 Weyl double copy in 4d
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Full four-gluon tree amplitude in Yang-Mills theory



Figure 3: The three Feynman diagrams corresponding to the s, t and u channels.

$$i\mathcal{A}_4^{\text{tree}} = g^2 \left( \frac{c_s n_s}{s} + \frac{c_t n_t}{t} + \frac{c_u n_u}{u} \right) \tag{1}$$

where

$$s = (p_1 + p_2)^2, \quad t = (p_2 + p_3)^2, \quad u = (p_3 + p_1)^2,$$
 (2)

are kinematic invariants of a four-point scattering process.

• The s-channel color factor

$$c_s = -2f^{a_1 a_2 b} f^{b a_3 a_4} \tag{3}$$

The s-channel kinematic numerator

$$n_{s} = -\frac{1}{2} \left[ (\epsilon_{1} \cdot \epsilon_{2}) p_{1}^{\mu} + 2(\epsilon_{1} \cdot p_{2}) \epsilon_{2}^{\mu} - (1 \leftrightarrow 2) \right] \cdot \left[ (\epsilon_{3} \cdot \epsilon_{4}) p_{3\mu} + 2(\epsilon_{3} \cdot p_{4}) \epsilon_{4\mu} - (3 \leftrightarrow 4) \right] - \frac{1}{2} s \left[ (\epsilon_{1} \cdot \epsilon_{3}) (\epsilon_{2} \cdot \epsilon_{4}) - (\epsilon_{1} \cdot \epsilon_{4}) (\epsilon_{2} \cdot \epsilon_{3}) \right]$$

$$(4)$$

The color/kinematic duality

$$c_s + c_t + c_u = 0, \quad n_s + n_t + n_u = 0$$
 (5)

• The kinematic factors satisfy the same relations as the color factors suggesting that they are mutually exchangeable.

[Bern, Carrasco, Johansson, 2008; 2010]

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• Swapping color factors for kinematic factors in the YM four-point amplitude, which gives a new gauge-invariant object that, is a four-graviton amplitude,

$$i\mathcal{A}_4^{\text{tree}}|_{g \to \frac{\kappa}{2}, c_i \to n_i} = i\mathcal{M}_4^{\text{tree}} = \left(\frac{\kappa}{2}\right)^2 \left(\frac{n_s^2}{s} + \frac{n_t^2}{t} + \frac{n_u^2}{u}\right) \tag{6}$$

- The new amplitude has the following properties:
  - the external states are captured by symmetric polarization tensors
  - the interactions are of the two-derivative type
  - invariant under linearized diffeomorphism transformations
- This amplitude should describe the scattering of four gravitons of Einstein theory.

# Double copy of classical solutions in the Kerr-Schild form

• The Kerr-Schild form is made by a null vector field  $k_{\mu}$  and a scalar function  $\phi$  in the form

$$g_{\mu\nu} = \eta_{\mu\nu} + \phi k_{\mu} k_{\nu} . \tag{7}$$

- One can show that  $k^{\mu}$  is null when contracted with both the full metric  $g_{\mu\nu}$  and the Minkowski one  $\eta_{\mu\nu}$ .
- The inverse metric is just

$$g^{\mu\nu} = \eta^{\mu\nu} - \phi k^{\mu} k^{\nu} . \tag{8}$$

• We can raise the index of k using both  $g^{\mu\nu}$  and  $\eta^{\mu\nu}$ .

# Stationary phase space

• In the stationary case in which all time derivatives vanish and  $k^0 = 1$  where all dynamics in the zeroth component contained in the scalar field, the Ricci tensor is reduced to

$$R^0{}_0 = \frac{1}{2}\nabla^2\phi \tag{9}$$

$$R^{i}_{0} = -\frac{1}{2}\partial_{j}\left[\partial^{i}(\phi k^{j}) - \partial^{j}(\phi k^{i})\right]$$
(10)

$$R^{i}_{\ j} = \frac{1}{2} \partial_{I} \left[ \partial^{i} (\phi k^{\prime} k_{j}) + \partial_{j} (\phi k^{\prime} k^{j}) - \partial^{\prime} (\phi k^{i} k_{l}) \right]$$
(11)

- Define a vector field  $A_{\mu} = \phi k_{\mu}$ , with the standard definition of field strength.
- The vacuum Einstein equation implies that the gauge field satisfies Maxwell equations

$$\partial_{\mu}F^{\mu\nu} = \partial_{\mu}\left[\partial^{\mu}(\phi k^{\nu}) - \partial^{\nu}(\phi k^{\mu})\right] = 0.$$
(12)

[Monteiro, O'Connell, White, 2014]

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• Using a pair of lightcone coordinates (*u*, *v*, *x*, *y*), we can express a pp-wave using

$$k_{\mu}\mathsf{d}x^{\mu} = \mathsf{d}u = \mathsf{d}z - \mathsf{d}t, \quad \phi = \phi(u, x, y)$$
 (13)

Einstein equation yields

$$\partial_{\mu}\partial^{\mu}\phi = 0. \tag{14}$$

• The gauge field  $A_{\mu} = \phi k_{\mu}$  a solution of the Maxwell equations.

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### Weyl double copy for type D solutions

• The Weyl double copy (WDC) formula

$$C_{ABCD} = \frac{1}{S} f_{(AB} f_{cD)}.$$
 (15)

• C<sub>ABCD</sub> is the anti-delf-dual part of the Weyl tensor,

$$C_{ABCD} = \frac{1}{4} \sigma^{\mu\nu}_{AB} \sigma^{\alpha\beta}_{CD} W_{\mu\nu\alpha\beta}.$$
 (16)

• f<sub>AB</sub> is defined from the field strength of the Maxwell tensor

$$f_{AB} = \frac{1}{2} \sigma^{\mu\nu}_{AB} F_{\mu\nu}.$$
 (17)

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[Luna, Monteiro, NIcholson, O'Connell, 2018]

Van der Waerden matrices

$$\sigma_{AB}^{\mu\nu} = \sigma_{A\dot{A}}^{[\mu} \tilde{\sigma}^{\nu]\dot{A}C} \epsilon_{CB}, \qquad \sigma_{A\dot{A}}^{\mu} = e_a^{\mu} \sigma_{A\dot{A}}^{a}, \qquad \sigma_{A\dot{A}}^{a} = \frac{1}{\sqrt{2}} (1, \sigma^i).$$

# Plebanski-Demianski metric

• The line-element in  $(p, q, \tau, \sigma)$  coordinates

$$ds^{2} = \frac{1}{(1-pq)^{2}} \left[ \frac{p^{2}+q^{2}}{P(p)} dp^{2} + \frac{p^{2}+q^{2}}{Q(q)} dq^{2} + \frac{P(p)}{p^{2}+q^{2}} (d\tau + q^{2}d\sigma)^{2} - \frac{Q(q)}{p^{2}+q^{2}} (d\tau - p^{2}d\sigma)^{2} \right], \quad (19)$$

where

$$P(p) = \gamma(1 - p^4) + 2np - \epsilon p^2 + 2mp^3,$$
 (20)

$$Q(q) = \gamma(1-q^4) - 2mq + \epsilon q^2 - 2nq^3.$$
<sup>(21)</sup>

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## The Weyl double copy

The Weyl spinor

$$C_{ABCD}\xi^{A}\xi^{B}\xi^{C}\xi^{D} = -\frac{3i}{2}\frac{(m+in)(1-pq)^{3}[(\xi^{1})^{2}-(\xi^{2})^{2}]^{2}}{(p-iq)^{3}}, \quad (22)$$

where  $\xi^{A} = (\xi^{1}, \xi^{2})$  is an arbitrary spinor.

The field strength spinor

$$f_{AB}\xi^{A}\xi^{B} = -\frac{1}{2}\frac{(m+in)(1-pq)^{3}[(\xi^{1})^{2}-(\xi^{2})^{2}]}{(p-iq)^{2}}$$
(23)

The scalar field

$$S = \frac{i}{6} \frac{(m+in)(1-pq)^3}{(p-iq)}$$
(24)

satisfies the wave equation in the flat background.

# Weyl double copy for type N solutions

• In the Newman-Penrose formalism, the type N WDC reads

$$\Psi_4 = \frac{1}{5} (\Phi_2)^2. \tag{25}$$

• The Bianchi identities

$$D\Psi_4 = (\rho - 4\epsilon)\Psi_4, \qquad \delta\Psi_4 = (\tau - 4\beta)\Psi_4.$$
 (26)

• The Maxwell's equation for degenerate Maxwell fields

$$D\Phi_2 = (\rho - 2\epsilon)\Phi_2, \qquad \delta\Phi_2 = (\tau - 2\beta)\Phi_2.$$
 (27)

The scalar field

$$D \log S - \rho = 0, \qquad \delta \log S - \tau = 0,$$
 (28)

which are integrable  $(\delta D - D\delta)S = \delta \rho - D\tau$ .

• Wave equation on type N background

$$\Box S = 0. \tag{29}$$

[Godazgar<sup>2</sup>, Monteiro, Veiga, Pope, 2020; 2021]

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Decomposition of the Weyl spinor and the Maxwell spinor

$$C_{ABCD} = a_{(A}b_Bc_Cd_D), \qquad f_{AB} = a_{(A}b_B). \tag{30}$$

• The Weyl double copy in the Newman-Penrose formalism

$$\Psi_{4} = 3c \frac{(\Phi_{2})^{2}}{S}, \qquad \Psi_{3} = 3c \frac{\Phi_{2}\Phi_{1}}{S},$$
  

$$\Psi_{2} = c \frac{2(\Phi_{1})^{2} + \Phi_{0}\Phi_{2}}{S}, \qquad (31)$$
  

$$\Psi_{1} = 3c \frac{\Phi_{0}\Phi_{1}}{S}, \qquad \Psi_{0} = 3c \frac{(\Phi_{0})^{2}}{S}.$$

[Godazgar<sup>2</sup>, Monteiro, Veiga, Pope, 2021; Mao, Zhao, 2023]

#### 5d vielbein formalism

The vielbein bases

$$I = e_0 = e^1, \quad n = e_1 = e^0, \quad m_i = e_i = m^i = e^i, \quad i = 2, 3, 4,$$
 (32)

where l and n are null, and  $m_i$  are spacelike. The basis vectors satisfy the orthogonal and normalization conditions

$$l \cdot l = n \cdot n = l \cdot m^{i} = n \cdot m^{i} = 0, \qquad (33)$$

$$l \cdot n = 1, \quad m^{i} \cdot m^{j} = \delta^{ij}. \tag{34}$$

Directional derivatives

$$D = I^{\mu}\partial_{\mu}, \qquad \Delta = n^{\mu}\partial_{\mu}, \qquad \delta_{i} = m_{i}^{\mu}\partial_{\mu}.$$
(35)

The following quantities are defined to denote the spin coefficients

$$L_{\mu\nu} = \nabla_{\nu} I_{\mu}, \quad N_{\mu\nu} = \nabla_{\nu} n_{\mu}, \quad M_{\mu\nu}^{k} = \nabla_{\nu} m_{\mu}^{k}.$$
(36)

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### 5d type N spacetime in CMPP classification

• Decomposition of the Weyl tensor

$$C_{\mu\nu\rho\sigma} = \underbrace{4C_{0i0j}n_{\{\mu}m_{\nu}^{i}n_{\rho}m_{\sigma}^{j}\}}_{\text{boost weight }+2} + \underbrace{8C_{010i}n_{\{\mu}l_{\nu}n_{\rho}m_{\sigma}^{i}\} + 4C_{0ijk}n_{\{\mu}m_{\nu}^{i}m_{\rho}^{j}m_{\sigma}^{k}\}}_{+1} + \underbrace{4C_{010i}n_{\{\mu}l_{\nu}n_{\rho}l_{\sigma}\} + 4C_{01ij}n_{\{\mu}l_{\nu}m_{\rho}^{i}m_{\sigma}^{j}\} + 8C_{0i1j}n_{\{\mu}m_{\nu}^{i}l_{\rho}m_{\sigma}^{j}\} + C_{ijkl}m_{\{\mu}^{i}m_{\nu}^{j}m_{\sigma}^{j}\}}_{0} + \underbrace{8C_{101i}l_{\{\mu}n_{\nu}l_{\rho}m_{\sigma}^{i}\} + 4C_{1ijk}l_{\{\mu}m_{\nu}^{i}m_{\rho}^{j}m_{\sigma}^{k}\}}_{-1} + \underbrace{4C_{1i1j}l_{\{\mu}m_{\nu}^{i}l_{\rho}m_{\sigma}^{j}\}}_{-2}.$$

For a type N solution, the non-zero Weyl scalars are  $\psi_{ij} = C_{\mu\nu\rho\sigma} n^{\mu} m^{i\nu} n^{\rho} m^{j\sigma}$  in the null frame

• Correspondingly,  $\phi_i = F_{\mu\nu} n^{\mu} m^{i\nu}$  are the non-zero Maxwell scalars of a degenerate Maxwell field [Coley, Milson, Pravda, Pravdova, 2004]

• Five  $\gamma$  matrices are chosen as the bases to connect the spacetime indices and spinor indices

$$\gamma^{\hat{\mu}}_{AB} = \begin{pmatrix} 0 & \sigma^{\hat{\mu}\alpha}{}_{\beta'} \\ \tilde{\sigma}^{\hat{\mu}}{}_{\alpha'}{}^{\beta} & 0 \end{pmatrix}, \quad \gamma^{4}_{AB} = -i \begin{pmatrix} \epsilon^{\alpha\beta} & 0 \\ 0 & \epsilon_{\alpha'\beta'} \end{pmatrix}.$$
(37)

•  $\hat{\mu} = 0, 1, 2, 3$  denotes the first four components of a 5d vector

- A, B = 1, 2 are spinor indices for 5d quantities
- $\alpha, \beta$  are now the spinor indices in the Van der Waerden matrices
- Any component of  $\gamma^{\mu}_{AB}$ , for any given number of A, B, and  $\mu = 0, 1, 2, 3, 4$ , defines a (2 × 2) matrix

[Monteiro, Nicholson, O'Connell, 2018]

Correspondingly, the spinor bases in 5d are chosen as

$$k_{1}^{A} = \begin{pmatrix} 0 \\ \bar{o}^{\alpha'} \end{pmatrix}, \quad k_{2}^{A} = \begin{pmatrix} o_{\alpha} \\ 0 \end{pmatrix}, \\ n_{1}^{A} = \begin{pmatrix} \iota_{\alpha} \\ 0 \end{pmatrix}, \quad n_{2}^{A} = -\begin{pmatrix} 0 \\ \bar{\iota}^{\alpha'} \end{pmatrix}.$$
(38)

One can simply package them as  $k_a^A = (k_1^A, k_2^A)$  and  $n_a^A = (n_1^A, n_2^A)$ , where indices *a* and *b* are packaged from two basis spinors.

# 5d WDC relation for type N spacetime

- The algebraic classification in the spinorial formalism is equivalent to the CMPP classification.
   [Monteiro, Nicholson, O'Connell, 2018]
- For the 5d type N spacetime, the non-zero little group bi-spinors for the Weyl tensor are

$$\psi_{abcd}^{(4)} = C_{\mu\nu\rho\sigma}\sigma^{\mu\nu}_{\ AB}\sigma^{\rho\sigma}_{\ CD}n^A_{\ a}n^B_{\ b}n^C_{\ c}n^D_{\ d}, \tag{39}$$

where  $\sigma^{\mu\nu}_{\ \ AB} = \frac{1}{2} (\gamma^{\mu}_{\ \ AC} \gamma^{\nu}{}^{C}_{\ \ B} - \gamma^{\nu}_{\ \ AC} \gamma^{\mu}{}^{C}_{\ \ B}).$ 

• A natural generalization of the WDC formula to 5d for the type N spacetime can be written as

$$\psi_{abcd}^{(4)} \propto \phi_{(ab}^{(2)} \phi_{cd}^{(2)},$$
 (40)

where  $\phi_{ab}^{(2)} = F_{\mu\nu}\sigma^{\mu\nu}_{\ AB}n^A_a n^B_b$  are the non-zero little group bi-spinors of a degenerate Maxwell tensor.

• More explicitly, we propose

$$\psi_{1111}^{(4)} = \frac{3c}{S_1} \phi_{11}^{(2)} \phi_{11}^{(2)}, \qquad \psi_{2222}^{(4)} = \frac{3c}{S_1'} \phi_{22}^{(2)} \phi_{22}^{(2)},$$
  

$$\psi_{1112}^{(4)} = \frac{3c}{S_2} \phi_{11}^{(2)} \phi_{12}^{(2)}, \qquad \psi_{2221}^{(4)} = \frac{3c}{S_2'} \phi_{22}^{(2)} \phi_{21}^{(2)}, \qquad (41)$$
  

$$\psi_{1212}^{(4)} = \frac{c}{S_3} [\phi_{11}^{(2)} \phi_{22}^{(2)} + 2(\phi_{12}^{(2)})^2].$$

We have introduced two independent complex scalar fields  $S_1$ ,  $S_2$  and one real scalar field  $S_3$ .

# 5D Weyl double copy relation in vielbein formalism

• Converting the WDC formulas (41) to the vielbein formalism

$$\psi_{22} - \psi_{33} - 2i\psi_{23} = \frac{3c}{S_1}[(\phi_2)^2 - (\phi_3)^2 - 2i\phi_2\phi_3],$$
  

$$\psi_{22} - \psi_{33} + 2i\psi_{23} = \frac{3c}{S_1'}[(\phi_2)^2 - (\phi_3)^2 + 2i\phi_2\phi_3],$$
  

$$\psi_{34} + i\psi_{24} = \frac{3c}{S_2}(\phi_3\phi_4 + i\phi_2\phi_4),$$
  

$$\psi_{34} - i\psi_{24} = \frac{3c}{S_2'}(\phi_3\phi_4 - i\phi_2\phi_4),$$
  

$$\psi_{44} = \frac{c}{S_3}[(\phi_2)^2 + (\phi_3)^2 + 4(\phi_4)^2].$$
  
(42)

The above equations present a natural extension of the WDC formulas to 5d for the type N spacetime.

#### Reductions of the type N spacetime

- We now consider a 4d-like reduction of the 5d type N spacetime where we impose the restrictions  $\psi_{4i} = 0$  for the Weyl scalars.
- The traceless property of the Weyl tensor then yields  $\psi_{22} = -\psi_{33}$ . Correspondingly, the Maxwell scalars and zeroth copies should be  $\phi_4 = 0, S_3 \rightarrow \infty$ , where the divergent scalar field  $S_3$  is chosen to turn off the extra component from the doubling in five dimensions.
- We introduce the following definitions

$$\begin{split} \Psi_{4} &= -(\psi_{22} + i\psi_{23}), \quad \bar{\Psi}_{4} = -(\psi_{22} - i\psi_{23}), \\ \varphi_{2} &= -\frac{1}{\sqrt{2}}(\phi_{2} + i\phi_{3}), \quad \bar{\varphi}_{2} = -\frac{1}{\sqrt{2}}(\phi_{2} - i\phi_{3}), \\ \delta &= \frac{1}{\sqrt{2}}(\delta_{2} - i\delta_{3}), \quad \bar{\delta} = \frac{1}{\sqrt{2}}(\delta_{2} + i\delta_{3}), \\ \hat{\delta} &= i\delta_{4}, \quad S = \frac{S_{1}}{c}. \end{split}$$
(43)

#### Bianchi identity and Maxwell's equation

$$D\Psi_{4} = -(L_{22} + iL_{23})\Psi_{4},$$
  

$$\delta\Psi_{4} = -\frac{1}{\sqrt{2}} \left[ (2L_{12} + 2M_{33}^{2} - L_{21}) + i(L_{31} - 2L_{13} + 2M_{32}^{2}) \right] \Psi_{4},$$
  

$$\hat{\delta}\Psi_{4} = \left[ (M_{42}^{3} - 2M_{24}^{3}) + i(L_{41} - 2L_{14} + M_{42}^{2}) \right] \Psi_{4},$$
(44)

and

$$D\varphi_{2} = -(L_{22} + iL_{23})\varphi_{2},$$
  

$$\delta\varphi_{2} = -\frac{1}{\sqrt{2}} [(L_{12} + M_{33}^{2} - L_{21}) + i(L_{31} - L_{13} + M_{32}^{2})]\varphi_{2},$$
  

$$\hat{\delta}\varphi_{2} = [(M_{42}^{3} - M_{24}^{3}) + i(L_{41} - L_{14} + M_{42}^{2})]\varphi_{2}.$$
(45)

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• The WDC formula in the reduced case is simply given by

$$\Psi_4 = \frac{1}{5} (\varphi_2)^2.$$
 (46)

- Now it is clear that the previously chosen divergent scalar field S<sub>3</sub> is to prevent the mixing term constructed from φ<sub>2</sub>φ<sub>2</sub> in the Weyl doubling.
- Such a term is not involved at all in the 4d WDC relation, and it is reasonable to have a special treatment in 5d by simply imposing that there is no Weyl scalar associated to the  $\varphi_2 \bar{\varphi}_2$  term.

# Integrability conditions and wave equation

• Substituting the WDC relation (46) into the Bianchi identity and simplifying them with the Maxwell's equation on the type N background, we obtain the equations for the scalar field.

$$D \log S = -(L_{22} + iL_{23}),$$
  

$$\delta \log S = \frac{1}{\sqrt{2}}(L_{21} - iL_{31}),$$
  

$$\hat{\delta} \log S = M_{42}^3 + i(M_{42}^2 + L_{41}).$$
(47)

- Following the treatment in 4d, we have proven the integrability conditions for the above differential equations.
- It is straightforward to verify that any solution of (47) solves the Klein-Gordon equation □S = 0 on the type N background.

#### New realization of 4d type D WDC

- We extend the type N analysis in 4d to the type D case.
- Any type D vacuum solution admits an algebraically general Maxwell scalar on the curved background that squares to give the Weyl scalar.
- Bianchi identities

$$D\Psi_2 = 3\rho\Psi_2, \quad \Delta\Psi_2 = -3\mu\Psi_2, \quad \delta\Psi_2 = 3\tau\Psi_2, \quad \bar{\delta}\Psi_2 = -3\pi\Psi_2.$$
(48)

Maxwell's equations

$$D\Phi_1 = 2\rho\Phi_1, \quad \Delta\Phi_1 = -2\mu\Phi_1, \quad \delta\Phi_1 = 2\tau\Phi_1, \quad \bar{\delta}\Phi_1 = -2\pi\Phi_1.$$
(49)

The type D WDC relation

$$\Psi_2 = \frac{1}{S} \Phi_1^2. \tag{50}$$

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#### New realization of 4d type D WDC

• Equations for the scalar field S

$$D\log S = 
ho, \quad \Delta\log S = -\mu, \quad \delta\log S = au, \quad ar{\delta}\log S = -\pi.$$
 (51)

#### Integrability conditions

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$$(\delta D - D\delta) \log S = \delta \rho - D\tau, \quad (\bar{\delta}\Delta - \Delta\bar{\delta}) \log S = \Delta\pi - \bar{\delta}\mu, (\Delta D - D\Delta) \log S = \Delta\rho + D\mu, \quad (\bar{\delta}\delta - \delta\bar{\delta}) \log S = \bar{\delta}\tau + \delta\pi, \quad (52) (\bar{\delta}D - D\bar{\delta}) \log S = \bar{\delta}\rho + D\pi, \quad (\delta\Delta - \Delta\delta) \log S = -\delta\mu - \Delta\tau.$$

• The WDC relation defines a scalar field

$$(\Box + 2\Psi_2)S = 0. \tag{53}$$

- We extend the type D WDC to 5d for a special class of vacuum solutions satisfying  $C_{\hat{a}\hat{b}\hat{c}4} = 0$ , where the integrability conditions consist of 10 equations.
- The scalar equaiton

$$\left[\Box - 2\psi_2 + 2(L_{41})^2\right]S = 0.$$
(54)

• Specially chosen fifth basis  $m^4$  from  $C_{\hat{a}\hat{b}\hat{c}4} = 0$  with the deviation tensor

$$\nabla_{\mu}m_{\nu}^{4} = L_{41}(m_{\mu}^{2}m_{\nu}^{2} + m_{\mu}^{3}m_{\nu}^{3} + l_{\mu}n_{\nu} + n_{\mu}l_{\nu}).$$
 (55)

• The expansion L<sub>41</sub> is the only non-zero component in the deviation that affects the evolution of the scalar S along the fifth direction m<sup>4</sup>.

# Thanks for your attention!

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