

# A brief review of loop quantum black hole models

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PRD 105, 024069; [PRD 104, 126003](#); PRD 102, 041502(R), and [ArXiv: 2404.02796](#);

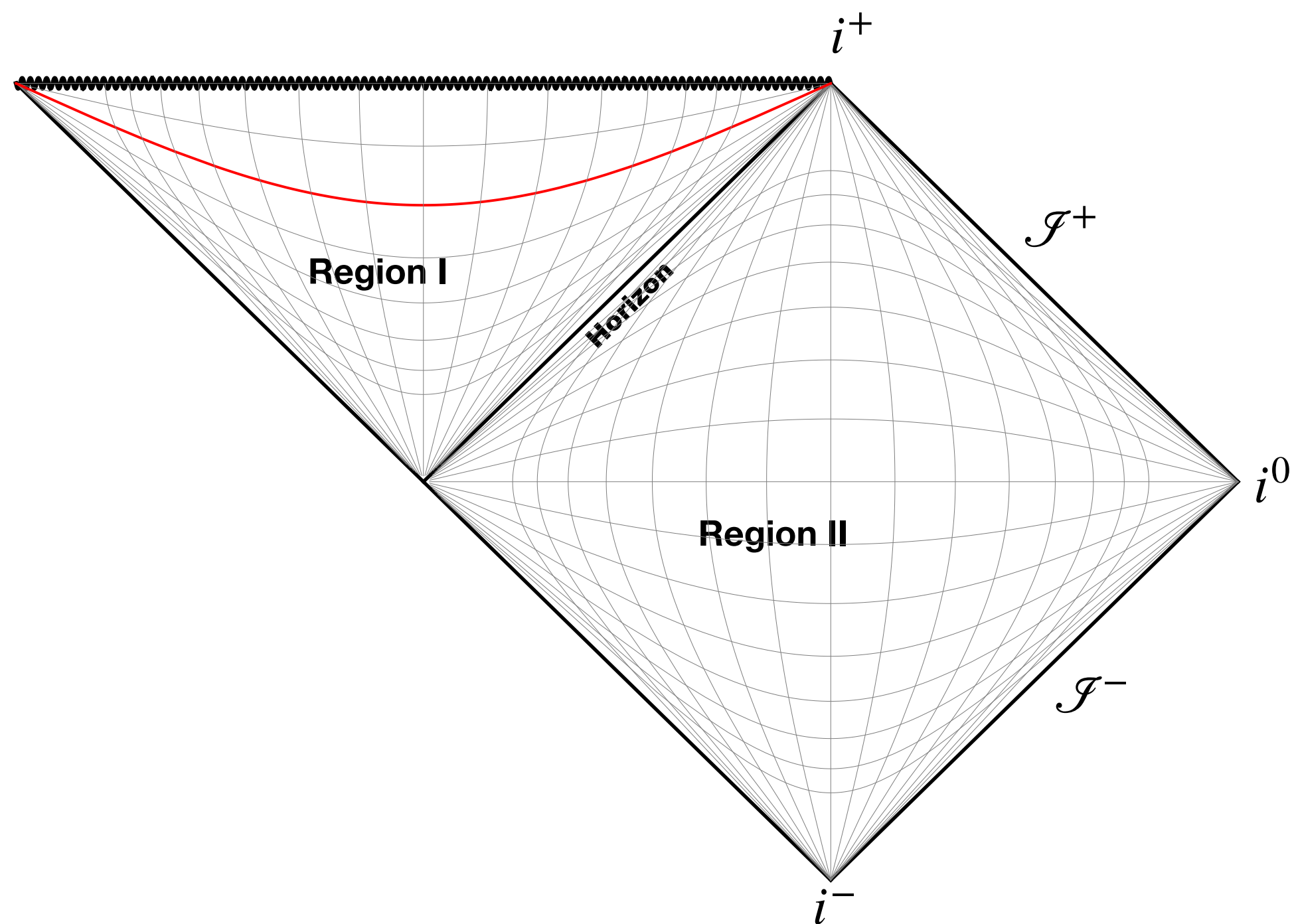
# Motivation and background

**The black holes of nature are the most perfect macroscopic objects there are in the Universe**

**—Subrahmanyan Chandrasekhar**

- **Black holes have been observed by gravitational wave and EHT; It sets the stage to probe quantum gravity;**
- **Spherically symmetric BH is a good candidate for QG research; It is neither as simple as cosmology nor very complicated.**
- **In Loop quantum gravity, our answer on quantum BH haven't form a unique picture; There are, e.g., Ashtekar-Bojowald paradigm [Ashtekar & Bojowald 05'], the SF qBH model [Rovelli, Haggard, Christodoulou, Speziale, Vilensky etc. 15', 16', Han, Qu, CZ 24'] and other models by the people in the community [Ashtekar, Bojowald, Bodendorfer, Boehmer, Chiou, Giesel, Gambini, Han, Liu, Modesto, Ma, Mehdi, Olmedo, Pullin, Singh, Vandersloot, Yang, Zhang and so on ]**

# Classical theories



Region I is a spacetime in its own right. The homogenous Cauchy slice has topology  $\mathbb{R} \times S^2$ :

- Standard Schwarzschild interior  $\iff$  EOM + general metric:

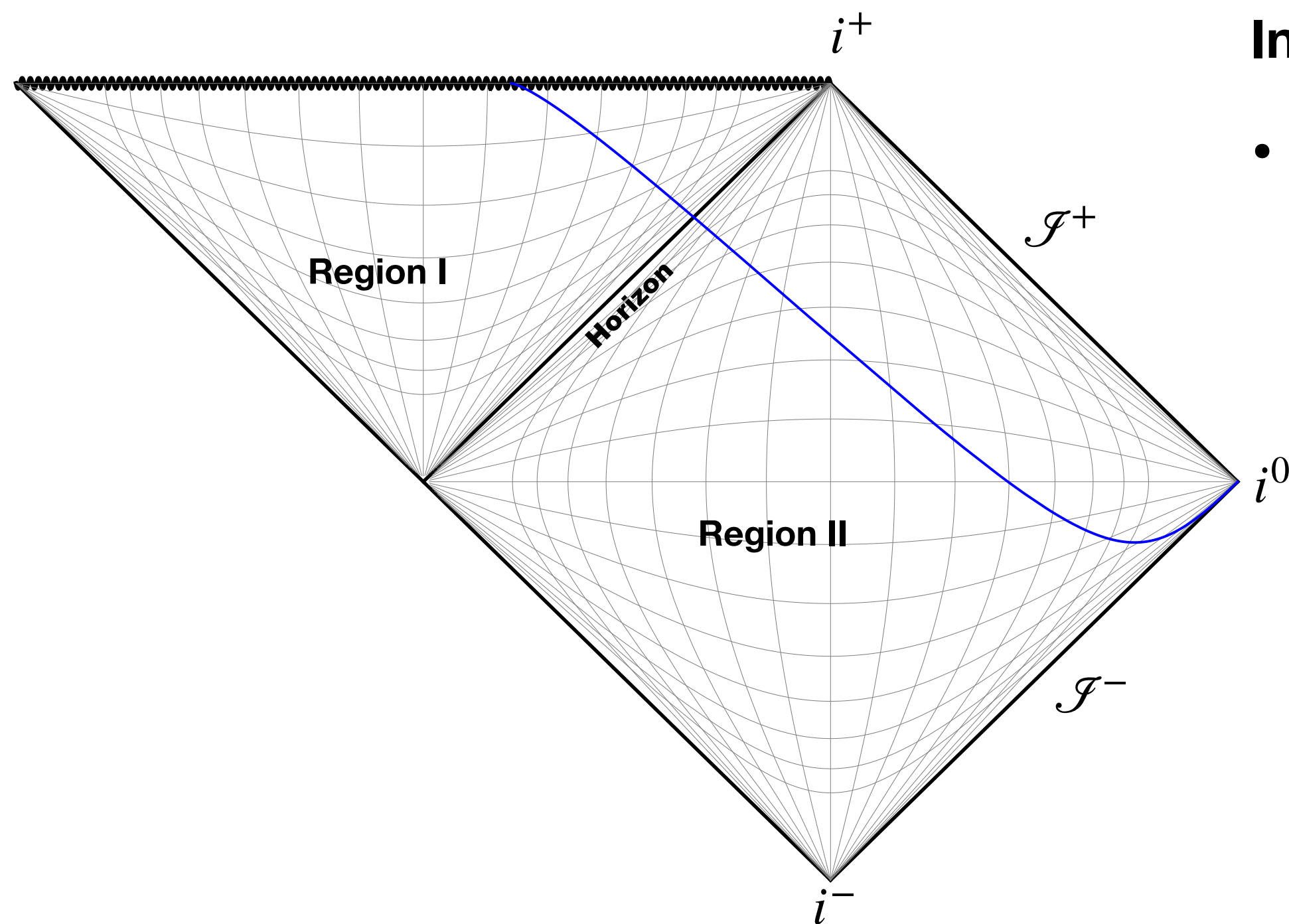
$$ds^2 = -N^2 dt^2 + \frac{p_b^2}{|p_c| L_0^2} dx^2 + |p_c| d\Omega^2$$

[Ashtekar, Olmedo & Singh 19', Zhang & CZ, 19']

This fact motives us to do (polymer) quantization on the surface  $\mathbb{R} \times S^2$  with symmetry  $\mathbb{T} \times SO(3)$

- Leads to a system with finite D.O.F,
- However, the exterior need further assumption.

# Classical theories



In the entire spacetime, the Cauchy surface have only the  $SO(3)$  symmetry

- Schwarzschild solution  $\Longleftarrow$  E.O.M + general metric

$$ds^2 = -N^2 dt^2 + \frac{(E^b)^2}{E^c} (dx + N^x dt)^2 + E^c d\Omega^2$$

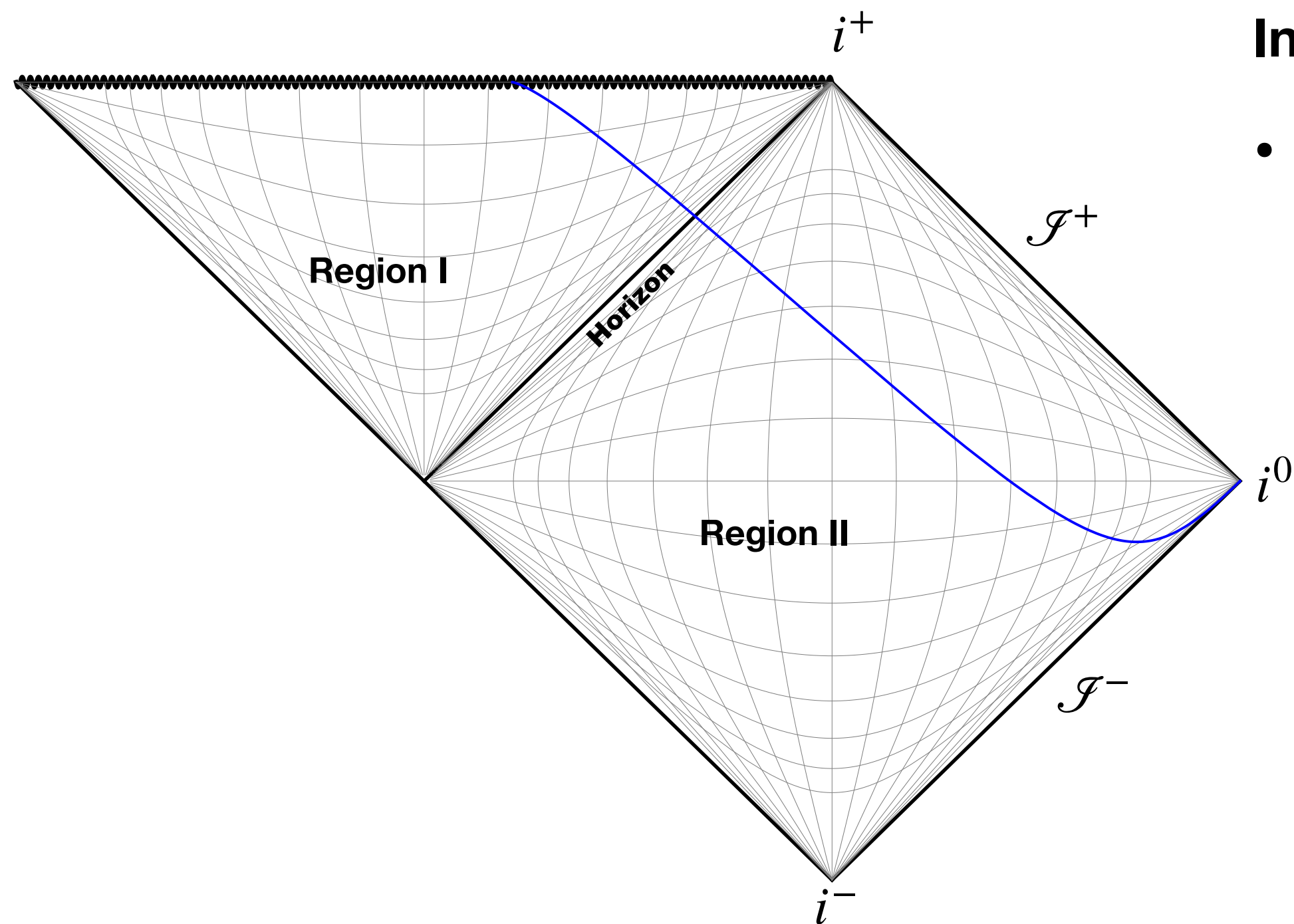
[Gambini, Olmedo & Pullin, 2015]

We are motived to do (polymer) quantization on the surface

$\mathbb{R} \times S^2$  with symmetry  $SO(3)$

- The entire spacetime is obtained,
- However, it leads to a system with infinity D.O.F.

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- The entire spacetime is obtained,
- However, it leads to a system with infinity D.O.F.

**It isn't guaranteed that the two quantum theories are the same even though their classical correspondences do so.**

# Quantum theories

$\mathbb{R} \times S^2$  with symmetry  $\mathbb{T} \times SO(3)$  :

$\mathbb{R} \times S^2$  with symmetry  $SO(3)$  :

**Graph:**

a vertex  $v$

**1-D lattice**

**Hilbert space basis:**

●  
 $|p_b, p_c\rangle$

● — ● — ● — ● — ●  
 $|E^b(v_1), E^c(v_1)\rangle \otimes |E^b(v_2), E^c(v_2)\rangle \otimes \dots \otimes |E^b(v_n), E^c(v_n)\rangle$

**Classical correspondence:**

$$E_i^a \tau^i \partial_a = p_c \tau_3 \sin \theta \partial_x + \frac{p_b}{L_0} \tau_2 \sin \theta \partial_\theta - \frac{p_b}{L_0} \tau_1 \partial_\phi$$

$$E_j^a \tau^j \frac{\partial}{\partial \sigma^a} = E^c(x) \sin(\theta) \tau_1 \partial_x + E^b(x) \tau_2 \sin(\theta) \partial_\theta + E^b(x) \tau_3 \partial_\phi$$

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**“Holonomy” operator**

$$\widehat{e^{i\lambda b}} |p_b, p_c\rangle = |p_b + \lambda, p_c\rangle, \quad \widehat{e^{i\lambda c}} |p_b, p_c\rangle = |p_b, p_c + \lambda\rangle$$

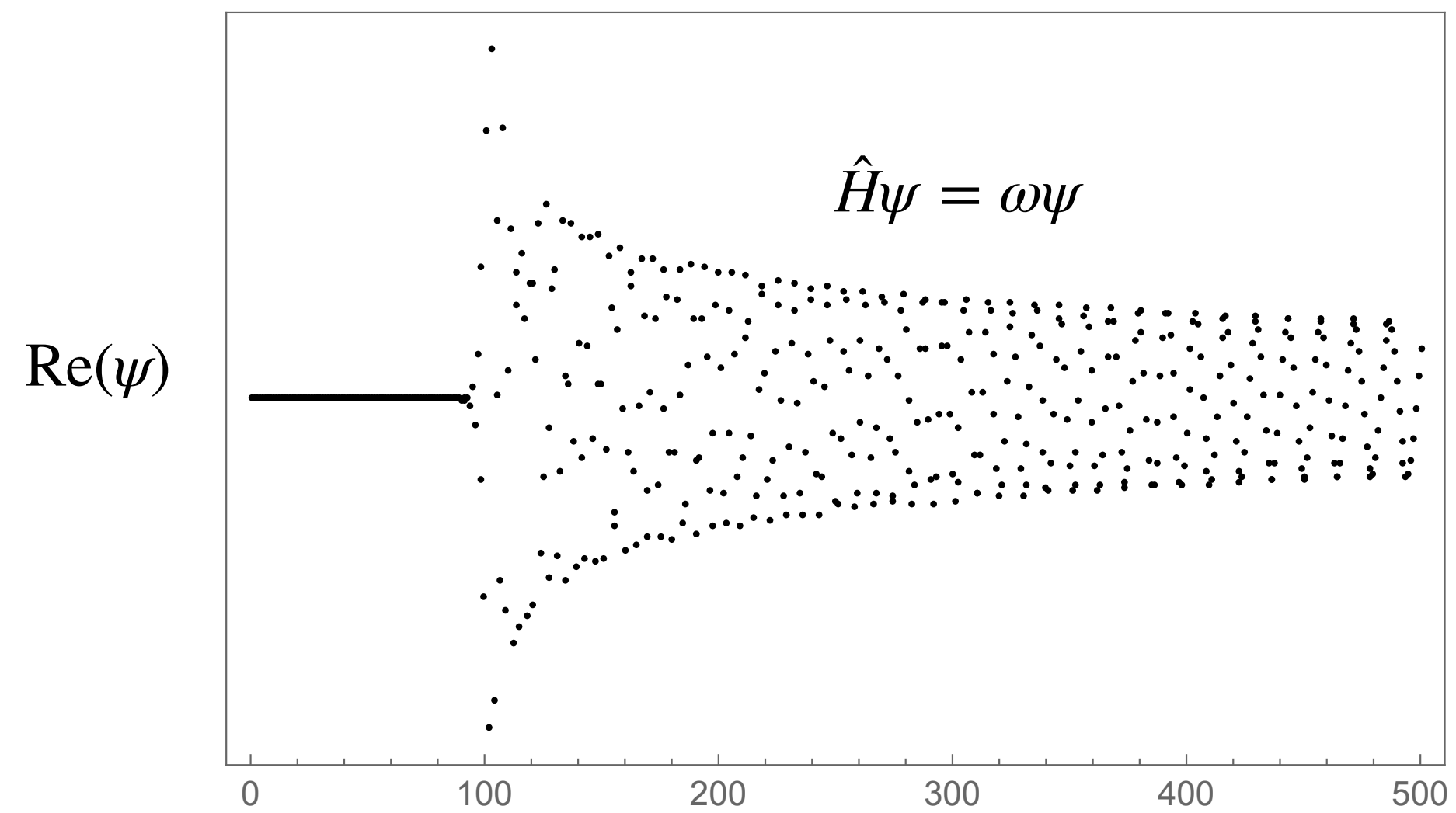
$$\begin{aligned} \widehat{e^{i\lambda(v)K_b(v)}} |E^b(v), E^c(v)\rangle &= |E^b(v) + \lambda(v), E^c(v)\rangle, \\ \widehat{e^{i\lambda(v)K_c(v)}} |E^b(v), E^c(v)\rangle &= |E^b(v), E^c(v) + \lambda(v)\rangle \end{aligned}$$

In contract to the Schordinger quantization, we have no differential operators like  $\hat{b}, \hat{c}, \hat{K}_b(v), \hat{K}_c(v)$ . However, the Hamiltonian involves the variables  $b, c, K_b(v), K_c(v)$ . We thus have to regularize them to be operators.

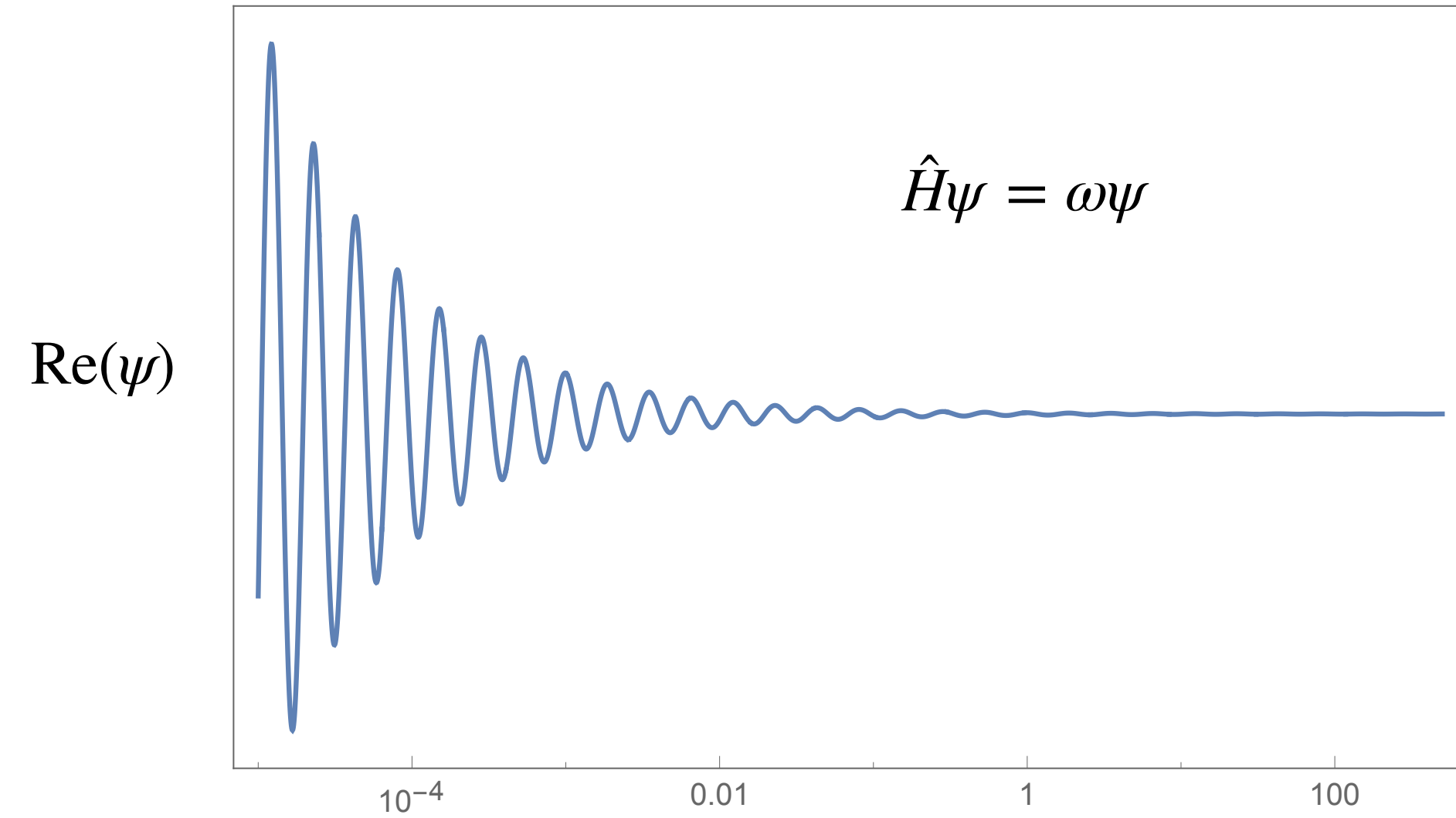
- Regularization leads to difference operator as the Hamiltonian, key to singularity resolution;

[CZ, Ma, Song, Zhang, 21', CZ, 21']

$$-i\frac{\partial}{\partial t}\psi = \hat{H}\psi, \quad \hat{H} = \sqrt{p}\hat{x}\sqrt{p}$$



$$(\hat{x}\psi)(p) \rightarrow i(\psi(p+1) - \psi(p-1))$$



$$\hat{x}\psi \rightarrow i\frac{d}{dp}\psi$$

[CZ, Lewandowski, Ma 19']



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- Regularization leads to difference operator as the Hamiltonian, key to singularity resolution;
- **Causes the ambiguities in LQBH.**

[CZ, Ma, Song, Zhang, 21', CZ, 21']

# Quantum theories

**The quantum dynamics of the homogeneous model:**

$$H[V] = -E_i^a E_j^b [\epsilon^{ij}_k F_{ab}^k - 2(1 + \gamma^2) K_{[a}^i K_{[a}^i] ] = 2p_b b c p_c + p_b^2 b^2 + \gamma^2 p_b^2$$

**Regularization gives** [\[Ashtekar, Bojowald 05', CZ, Ma, Song, Zhang 21'\]](#)

$$H[V]^{(\tilde{\delta}_b, \tilde{\delta}_c)} = 2p_b \frac{\sin(\tilde{\delta}_b b)}{\tilde{\delta}_b} p_c \frac{\sin(\tilde{\delta}_c c)}{\tilde{\delta}_c} + p_b^2 \frac{\sin^2(\tilde{\delta}_b b)}{\tilde{\delta}_b^2} + \gamma^2 p_b^2$$

**Classically,**  $H[V] = \lim_{\tilde{\delta}_b, \tilde{\delta}_c \rightarrow 0} H[V]^{(\tilde{\delta}_b, \tilde{\delta}_c)}$  **but in quantum theory,**  $\widehat{H[V]} = \lim_{\tilde{\delta}_b, \tilde{\delta}_c \rightarrow \delta_b, \delta_c} \widehat{H[V]}^{(\tilde{\delta}_b, \tilde{\delta}_c)}$

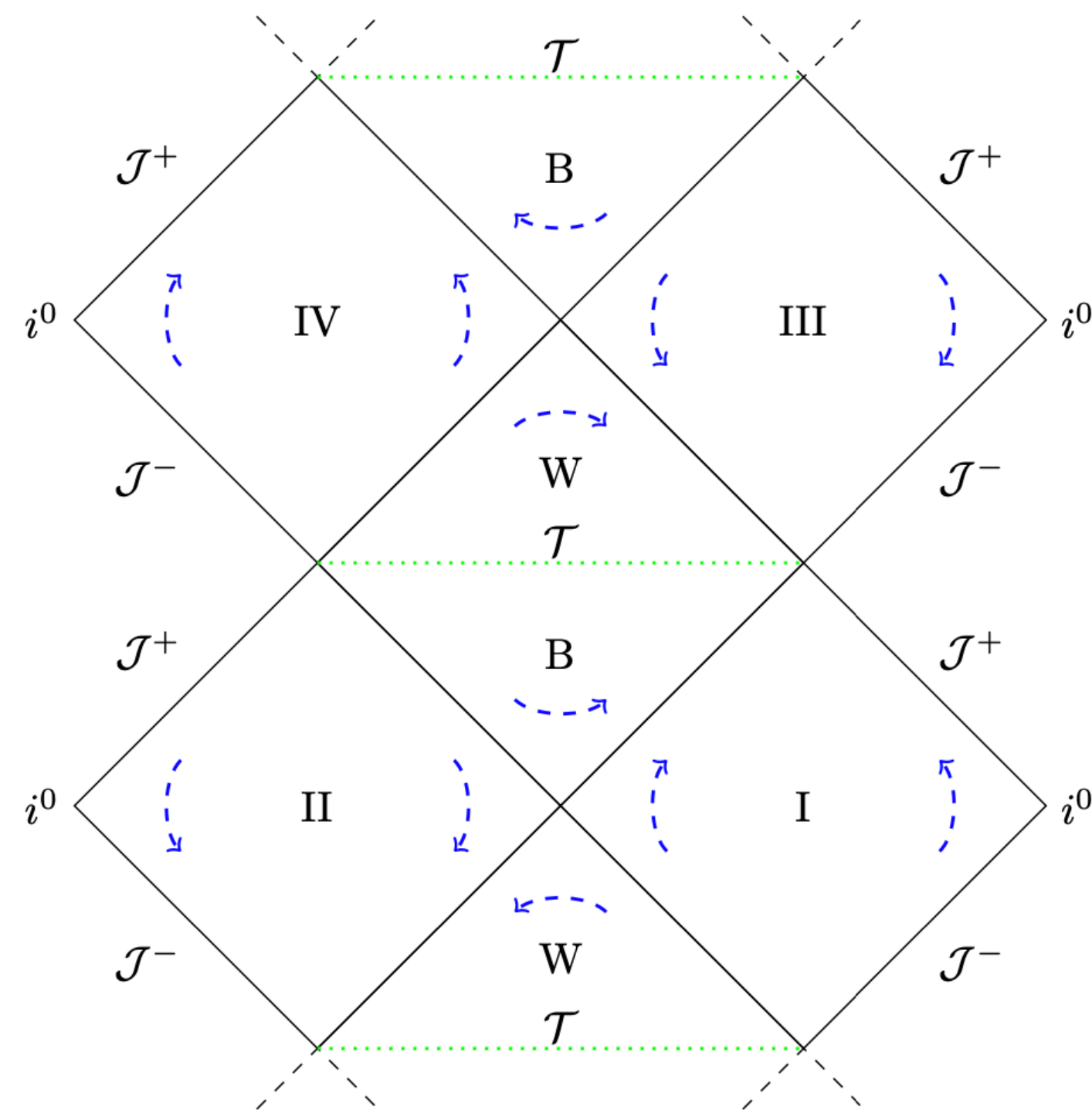
**Ambiguities arise due to various choices of  $\delta_b, \delta_c$  :**

- $\mu_0$ —scheme, constant  $\delta_b, \delta_c$  ; [\[Boehmer Vanderslhoot 07', Chiou 08'\]](#)
- $\bar{\mu}$ —scheme,  $\delta_b, \delta_c$  being phase space function; [\[Chiou 08'\]](#)
- **New scheme,  $\delta_b, \delta_c$  being function of dynamical trajectories.** [\[Corichi, Singh 16', Ashtekar, Olmedo Singh 18'\]](#)

# Quantum theories

## Some results:

- **Effective dynamics: singularity resolution, BH-WH transition, etc.**  
[Boehmer Vandershooft 07', Chiou 08', Corichi, Singh 16', Ashtekar, Olmedo Singh 18']
- **Quantum dynamics: discreteness of BH mass at the dynamical level;** [CZ, Ma, Song, Zhang 20' & 21']

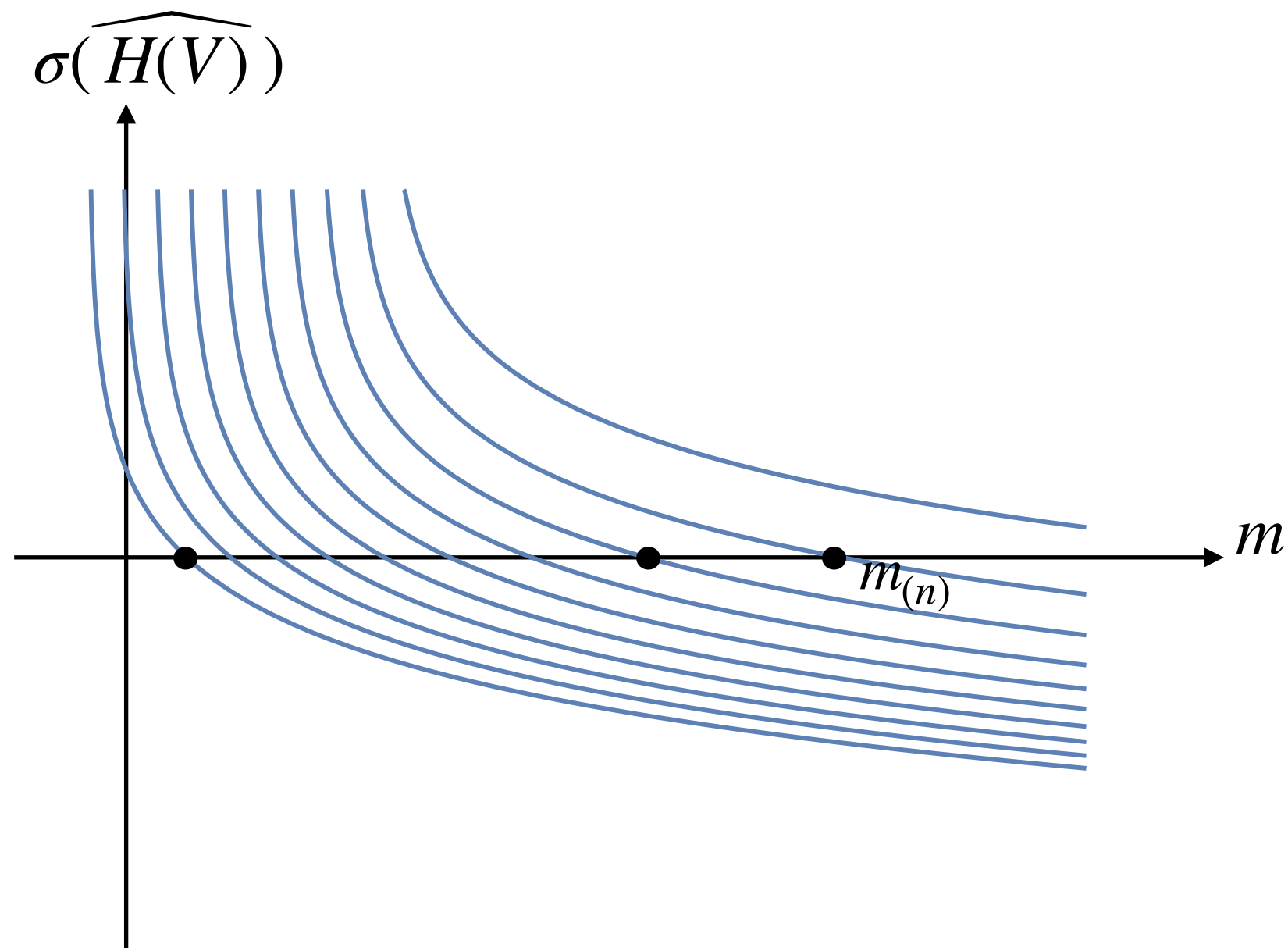


[Ashtekar, Olmedo Singh 18']

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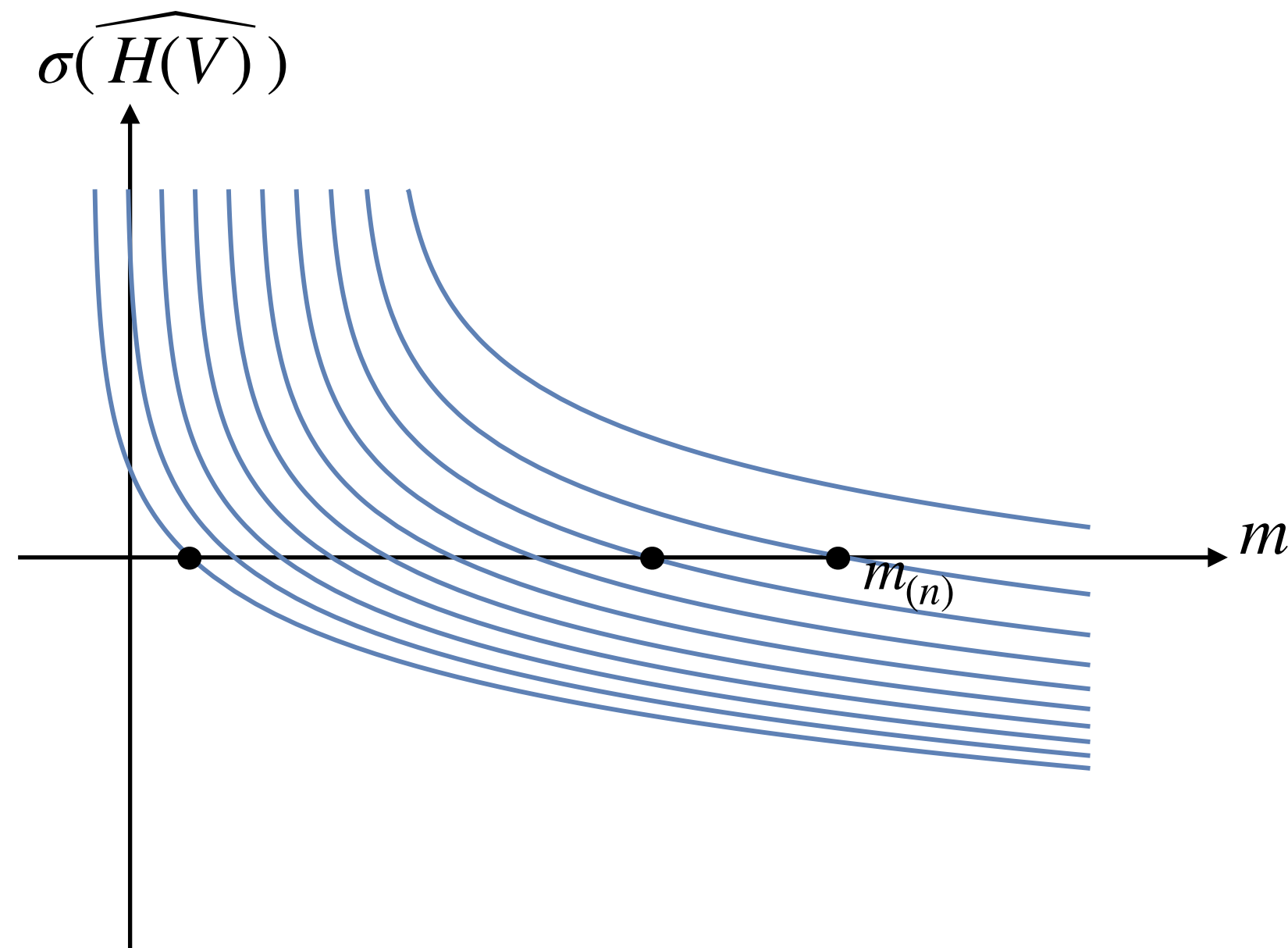
Constraint equation:  $\widehat{H[V]} = 0$ :

- Only for **countably many** values  $m_{(n)}$ , one can obtain solutions;
- The minimal value  $m_{(0)}$  is **not vanishing**.

# Quantum theories

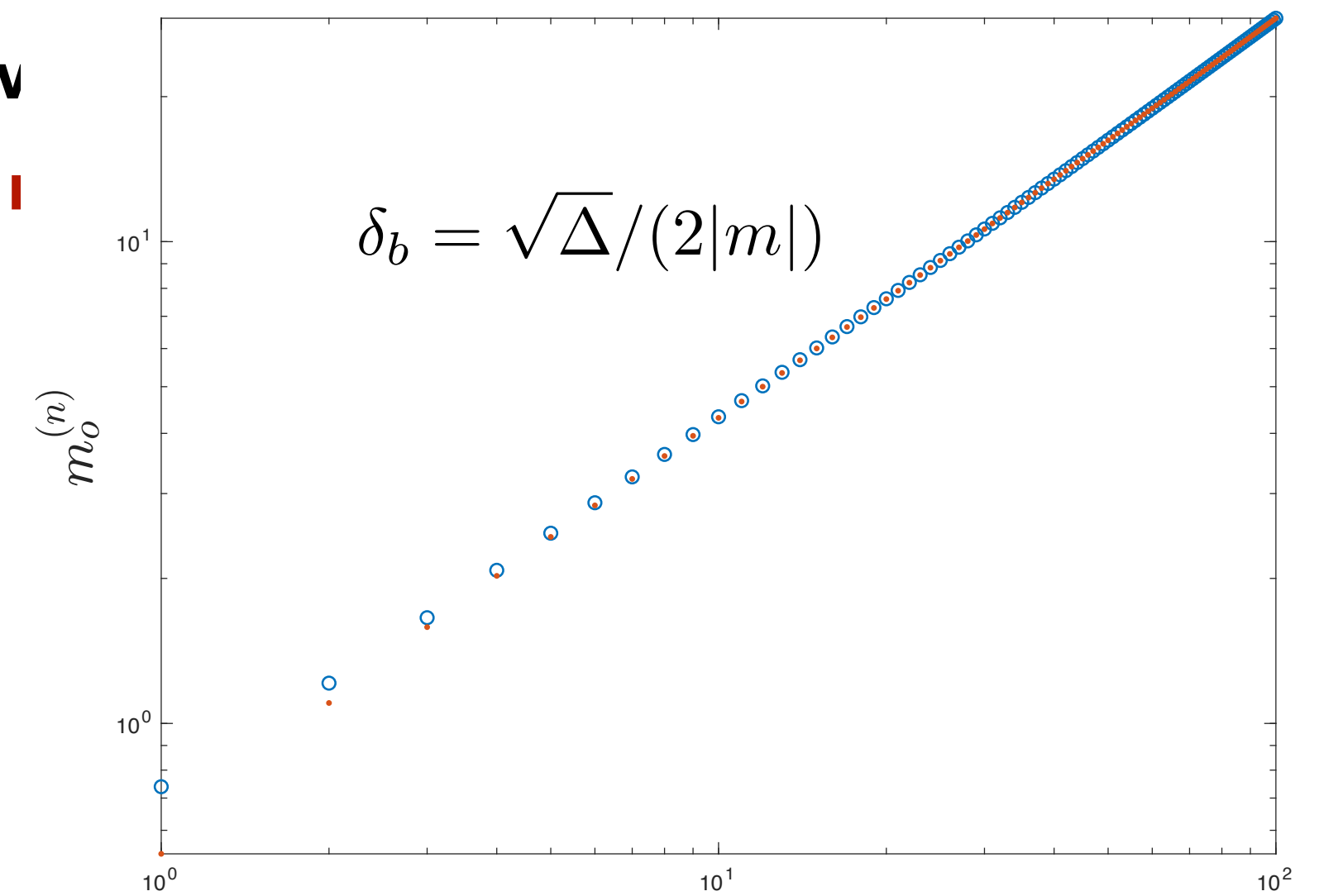
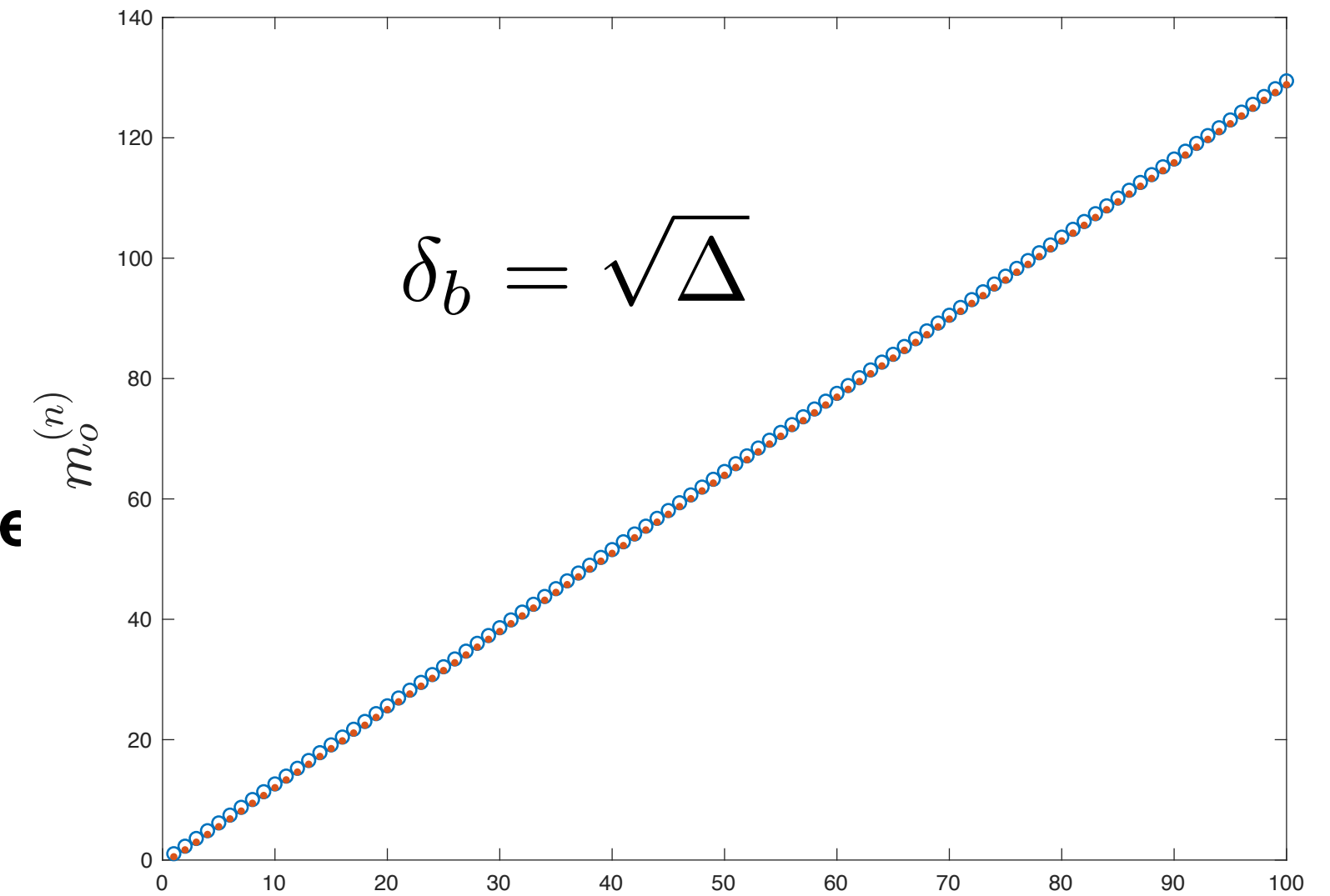
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- **Quantum dynamics:** discreteness of BH mass at the dynamical level



- Constraint equation:  $\widehat{H[V]}$
- Only for **countably many**  $v$
  - The minimal value  $m_{(0)}$  is **!**

**Not for  $\bar{\mu}$  scheme!**



# Quantum theories

The quantum dynamics of the non-homogeneous model:

$$C = \frac{-1}{|E^b| \sqrt{|E^c|}} \left( 8E^c E^b K_c K_b + 2E^{b2} K_b^2 + 2E^{b2} \right) - \frac{E^{c'2}}{2|E^b| \sqrt{|E^c|}} + \left( \frac{2E^c E^{c'}}{|E^b| \sqrt{|E^c|}} \right)'$$

$$\mathcal{V} = E^b K'_b - E^{c'} K_c$$

$$H_T = \int dx (N(x)C(x) + N^x(x)\mathcal{V}(c))$$

$\bar{\mu}$ -scheme is chosen to quantize the Hamiltonian constraint. However, there still are ambiguities caused by different approaches to deal with the constraint algebra:

- **Gambini-Olmedo-Pullin model:** choose phase space dependent lapse function and shift vector to modify the constraint algebra to get a “true” Lie algebra; [\[Gambini, Olmedo, Pullin 14' & 20\]](#)
- **Kelly-Santacruz-Wilson-Ewing model:** Introduce the Areal gauge to do gauge fixing; [\[Kelly, Santacruz, Wilson-Ewing 20'&22', see Giesel, Li Singhm, Weigl 21' for further discussion on gauge fixing\]](#)
- **Han-Liu model:** consider the phase-space reduce quantization where the constraints are solve at the classical level. [\[Han, Liu 20', CZ 21', see, e.g., Giesel, Tambornino, Thiemann 09' for phase-space reduced quantization\]](#)

Some results in the Han-Liu model: [CZ 21']

- The quantum dynamics is studied by path integral formulation,
- The effective dynamics is coincide with the heuristic one where one just replaces  $K_a, K_b$  in the Hamiltonian by sine functions and solve the Hamilton's equations,
- For small BH mass, the effective dynamics is different from the heuristics one.

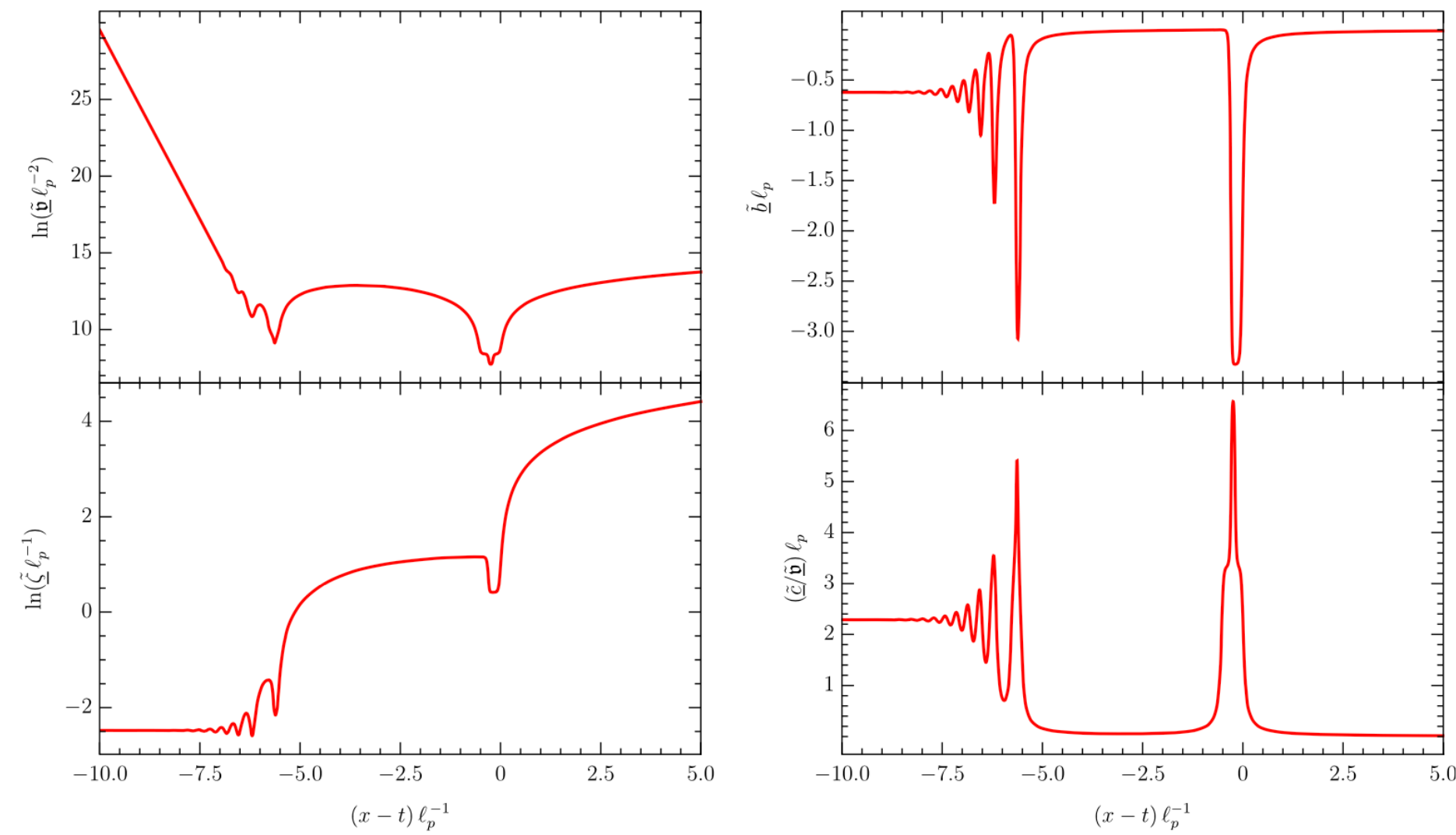


FIG. 1. Solutions to (6.20), (6.22), (6.23), and (6.24). The initial data are chosen such that  $\tilde{\mathfrak{u}}(y_0) = 6\pi F_0 y_0$ ,  $\tilde{\mathfrak{b}}(y_0) = -\frac{1}{6\pi y_0}$ ,  $\tilde{\xi}(y_0) = (\frac{3}{2}\sqrt{F_0}y_0)^{2/3}$ ,  $\tilde{\mathfrak{z}}(y_0) = \frac{3F_0}{2}$ , and  $\tilde{\xi}'(y_0) = \frac{2}{3}(\frac{3}{2}\sqrt{F_0}y_0)^{-1/3}$  with  $y_0 = 10^5\ell_p$  and  $F_0 = 10^4\ell_p$ . The initial data are choose by considering the Schwarzschild solution with  $2GM = F_0$  in Lemâitre coordinate at  $x - t = y_0$ . The parameters are set to be  $\Delta = 0.1$  and  $\beta = 0.2375$ .

The spacetime ends up in a Nariai spacetime  $ds_2 \times S_2$

[Han, Liu 20', CZ 21']

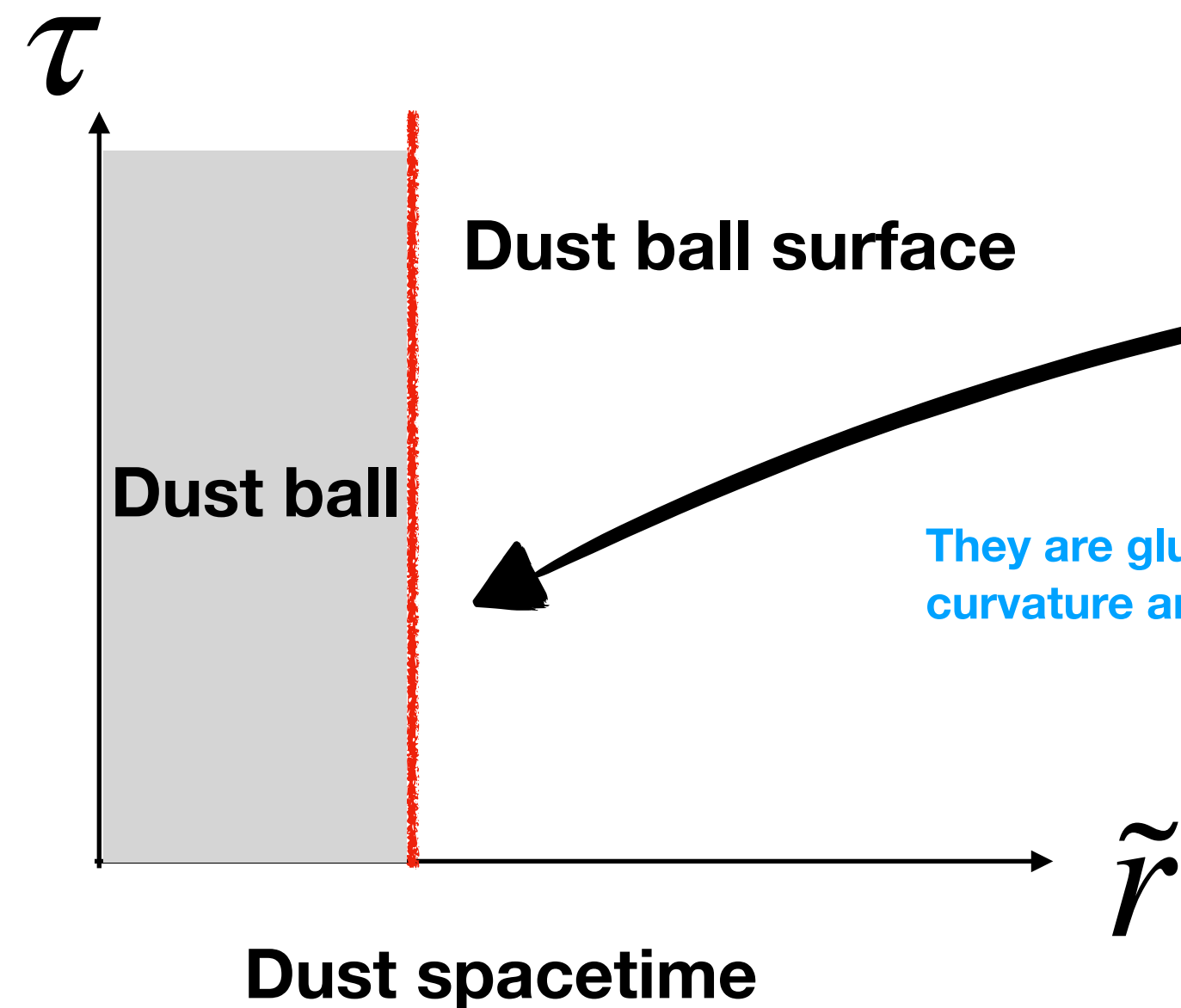
# Matter collapsing in LQG

Consider LQC dust to determine the exterior spacetime by junction condition. [Lewandowski, Ma, Yang, CZ 23']

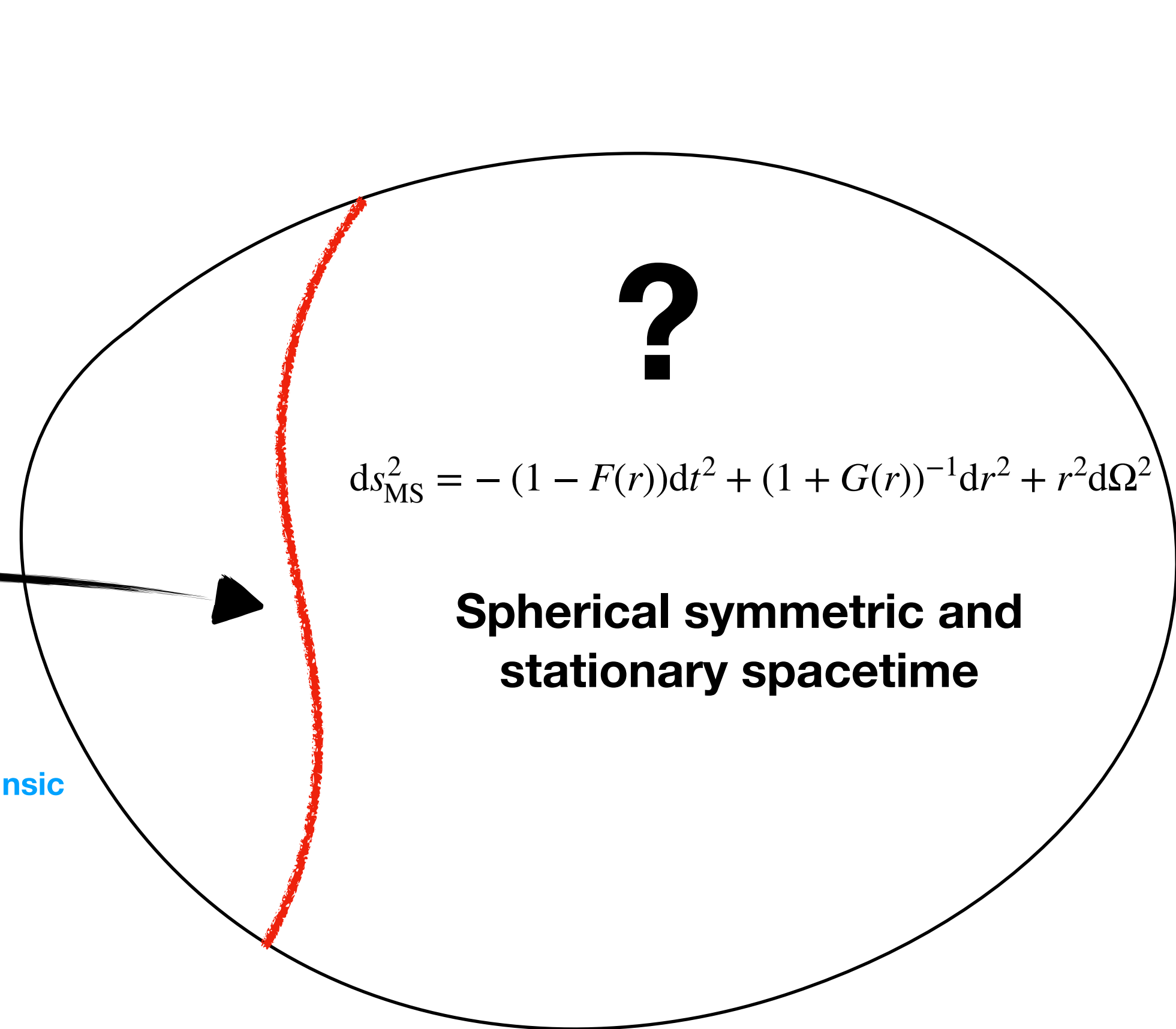
Gravity coupled to homogeneous pressureless dust

$$ds_{\text{APS}}^2 = -d\tau^2 + a(\tau)(d\tilde{r}^2 + \tilde{r}^2 d\Omega^2)$$

$$\text{Dynamics: } \mathbb{H}^2 = \frac{8\pi G}{3} \rho \left(1 - \frac{\rho}{\rho_c}\right), \quad \partial_\tau(\rho a^3) = 0$$



They are glued such that the reduced metric and extrinsic curvature are continuous



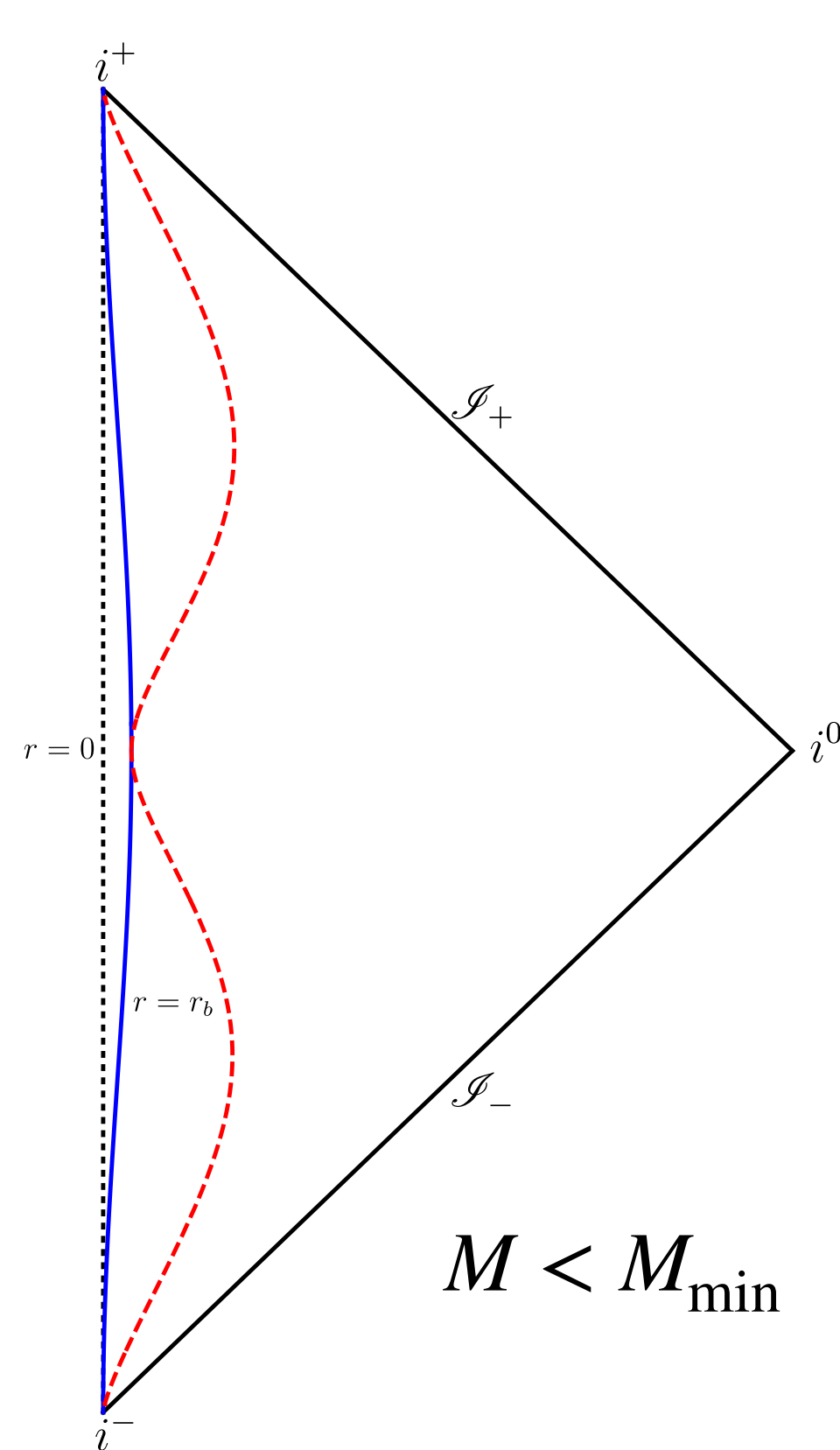


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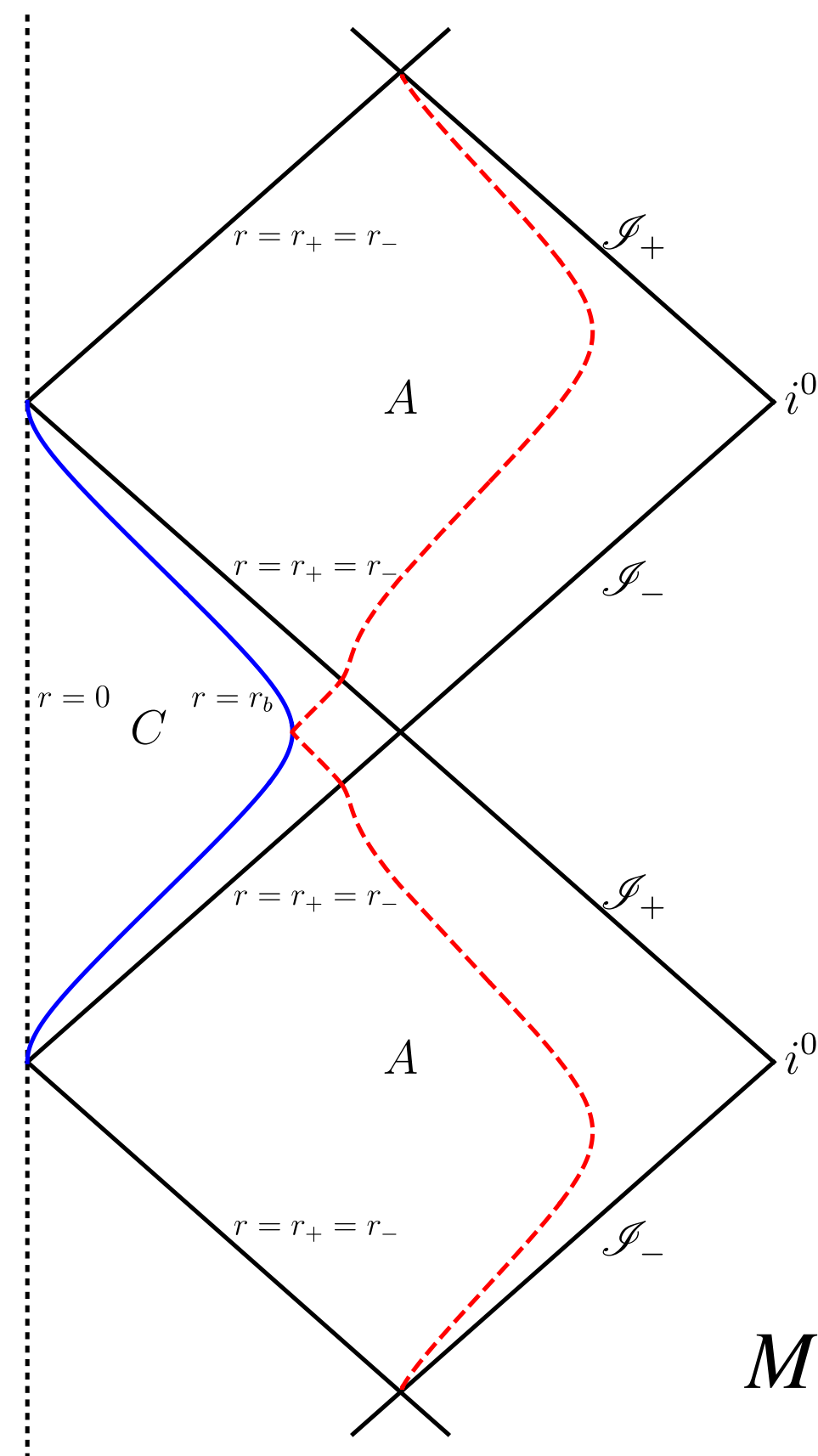
[Lewandowski, Ma, Yang, CZ 22']

$$ds_{\text{MS}}^2 = - \left( 1 - \frac{2GM}{r} + \frac{\alpha G^2 M^2}{r^4} \right) dt^2 + \left( 1 - \frac{2GM}{r} + \frac{\alpha G^2 M^2}{r^4} \right)^{-1} dr^2 + r^2 d\Omega^2, \quad \alpha = 16\sqrt{3}\pi\gamma^3 \ell_p^2$$

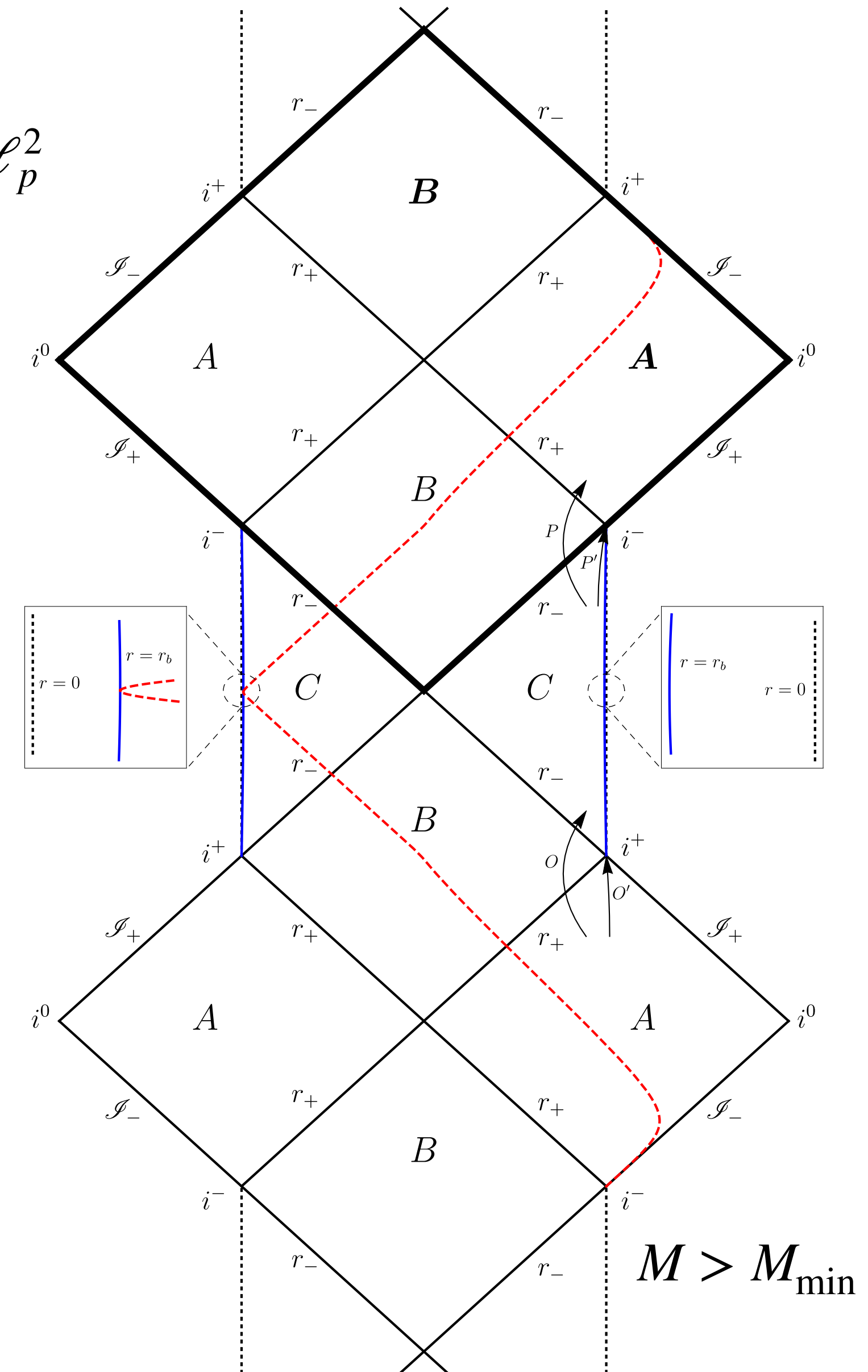
The result is the same as the one given by Kelly, Santacruz, Wilson-Ewing 20' but from loop quantization approach



$M < M_{\min}$



$M = M_{\min}$



$M > M_{\min}$

# Observing effect of LQBH

[Yang, CZ, Ma 23']

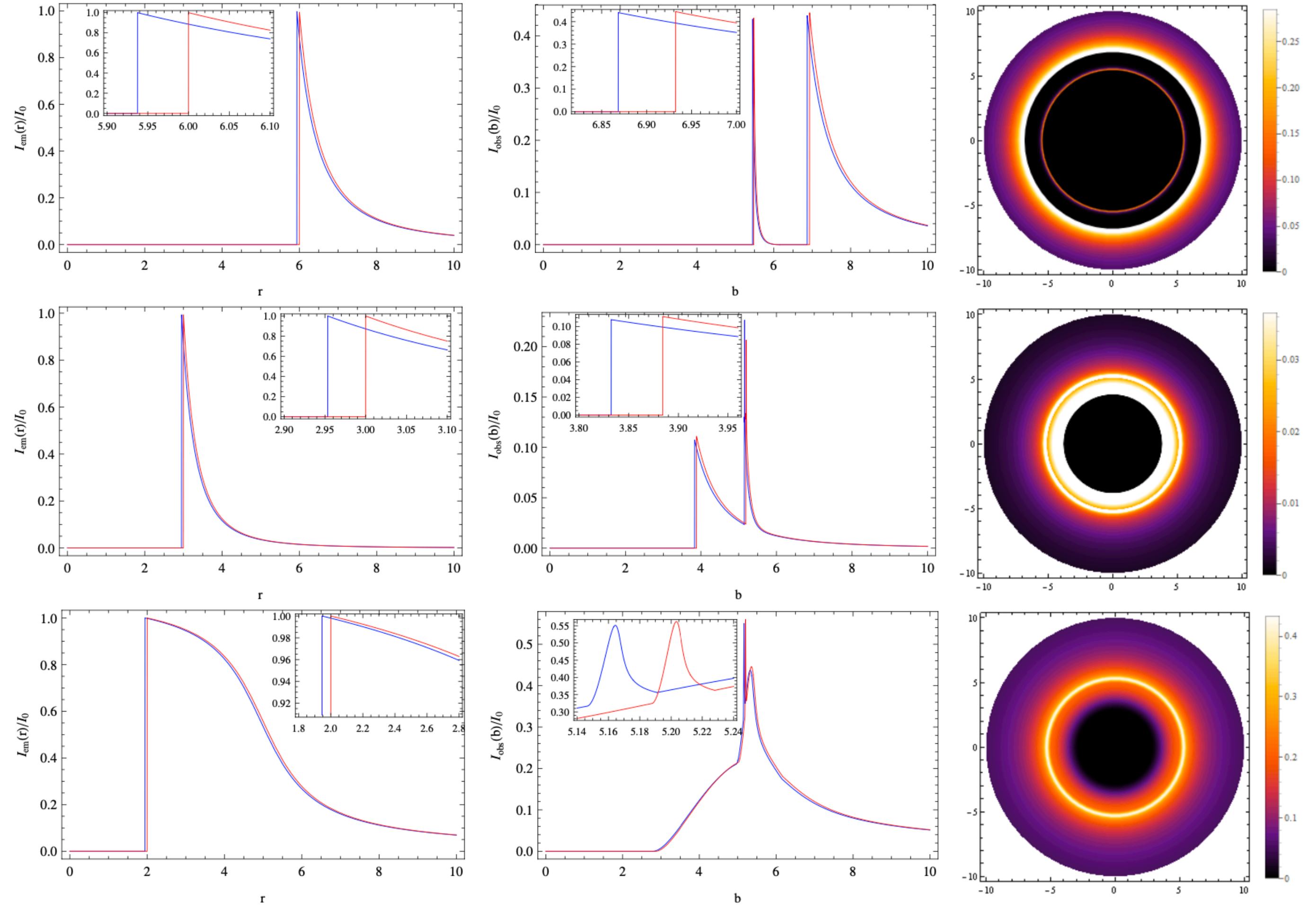
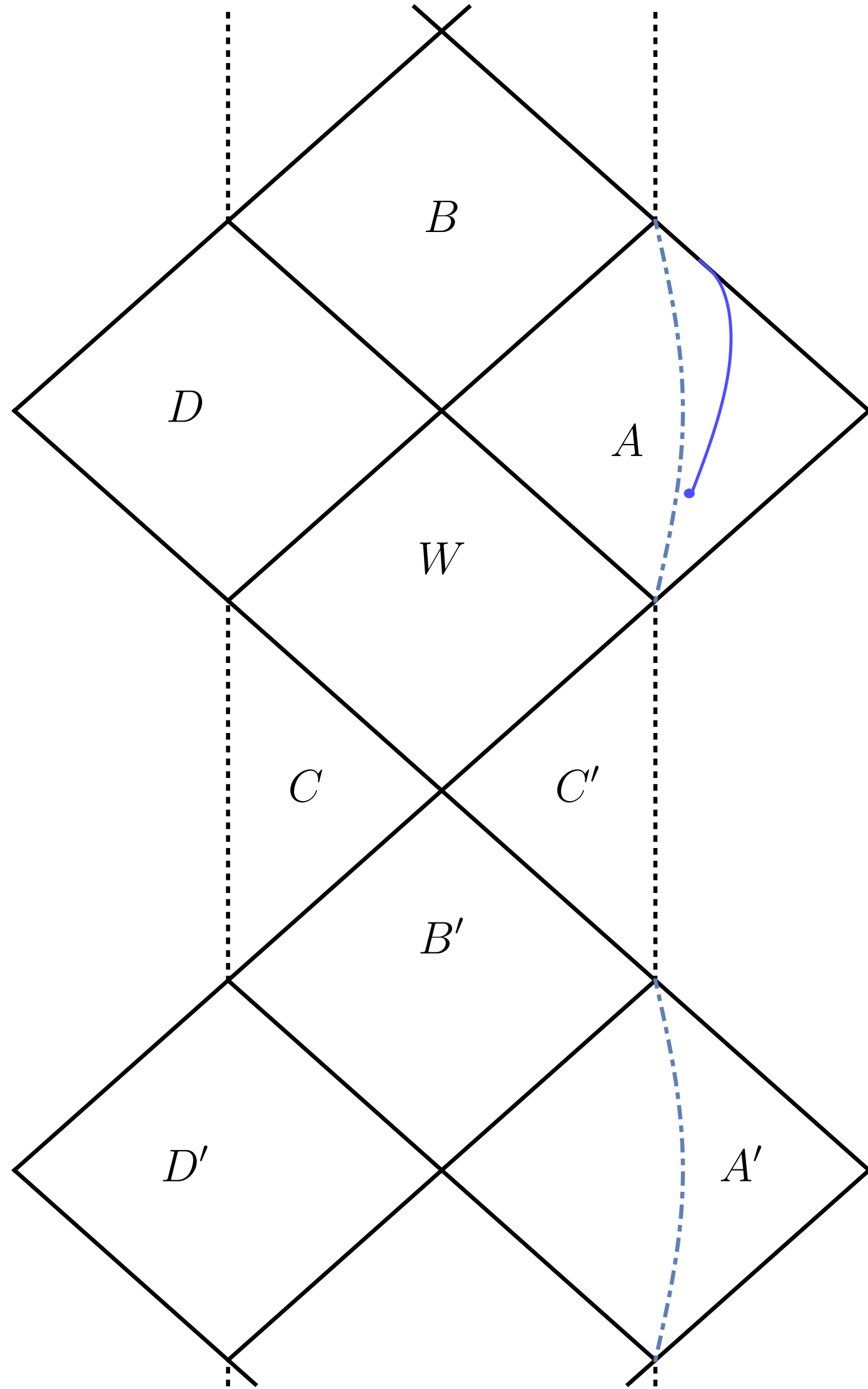
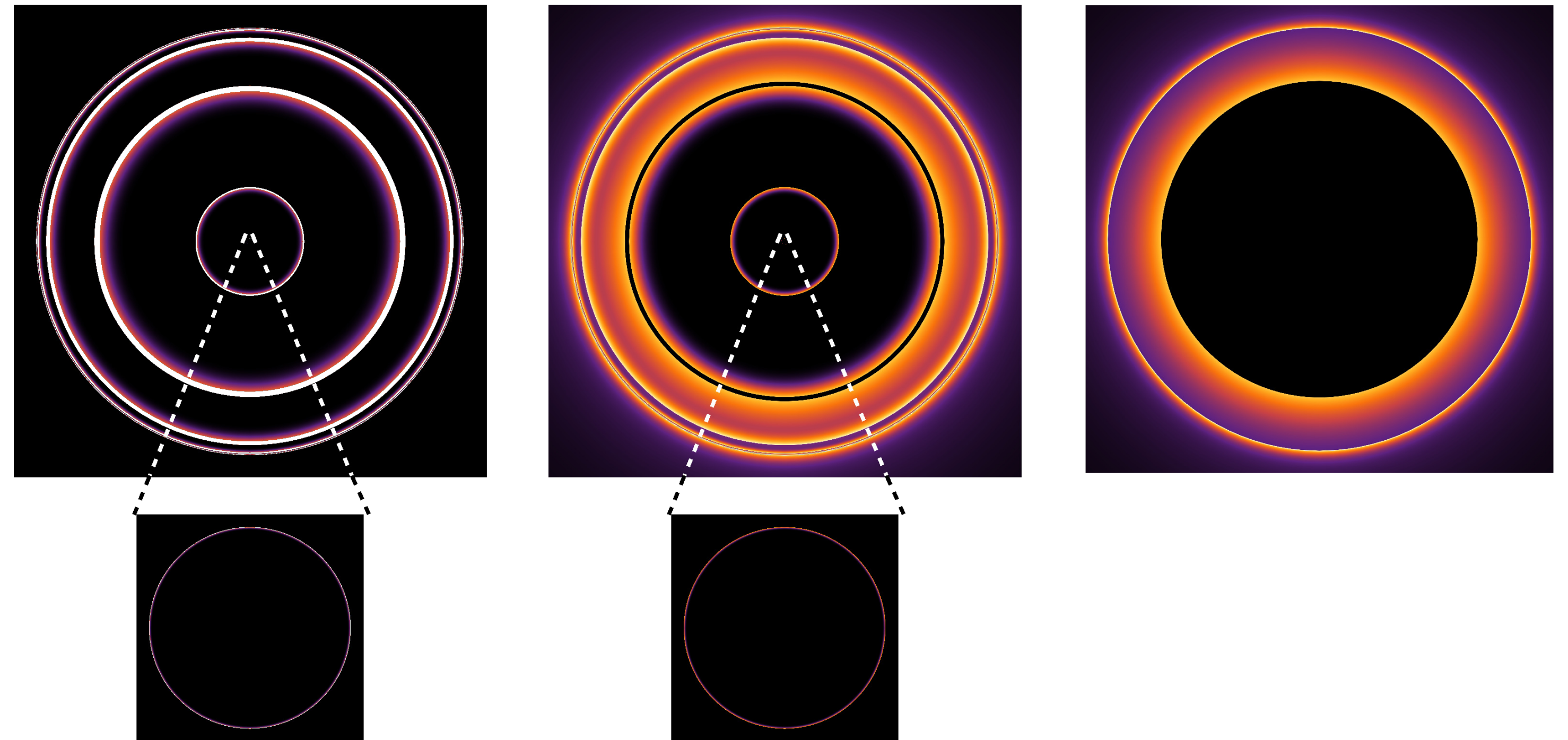
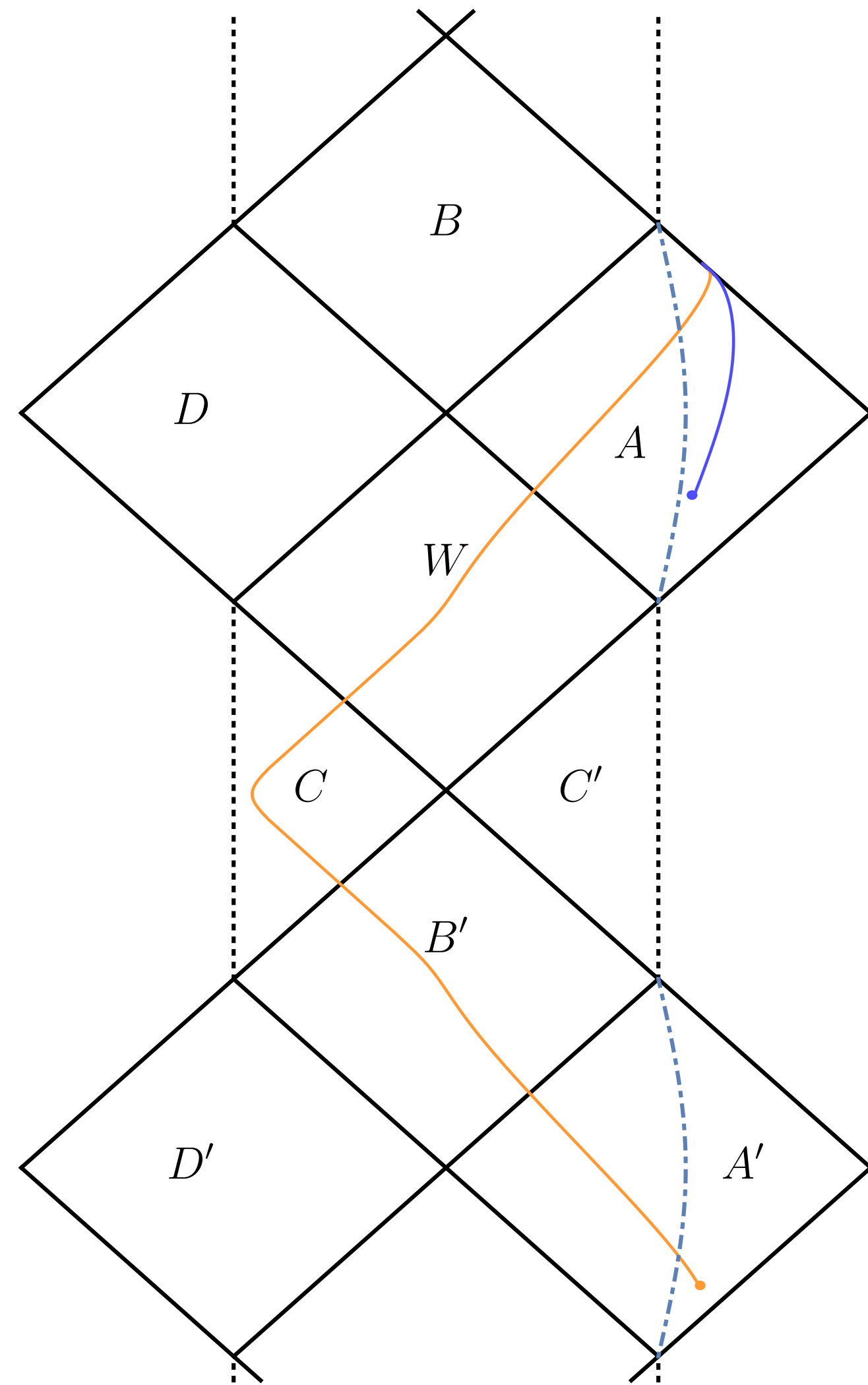


FIG. 6. The observational appearances of the thin disk near the BHs with the three different profiles. In each row, the first two panels show the emission intensity  $I_{\text{em}}/I_0$  and observational intensity  $I_{\text{obs}}/I_0$ , normalized to the maximum value  $I_0$ , of a thin disk near the quantum-corrected BH (blue) compared to those of the Schwarzschild BH (red), and the third panel depicts the density plot of  $I_{\text{obs}}/I_0$  of a thin disk near the quantum-corrected BH. The parameters are  $R_s = 2$ ,  $\gamma = 1$  and  $\Delta = 0.1$ .

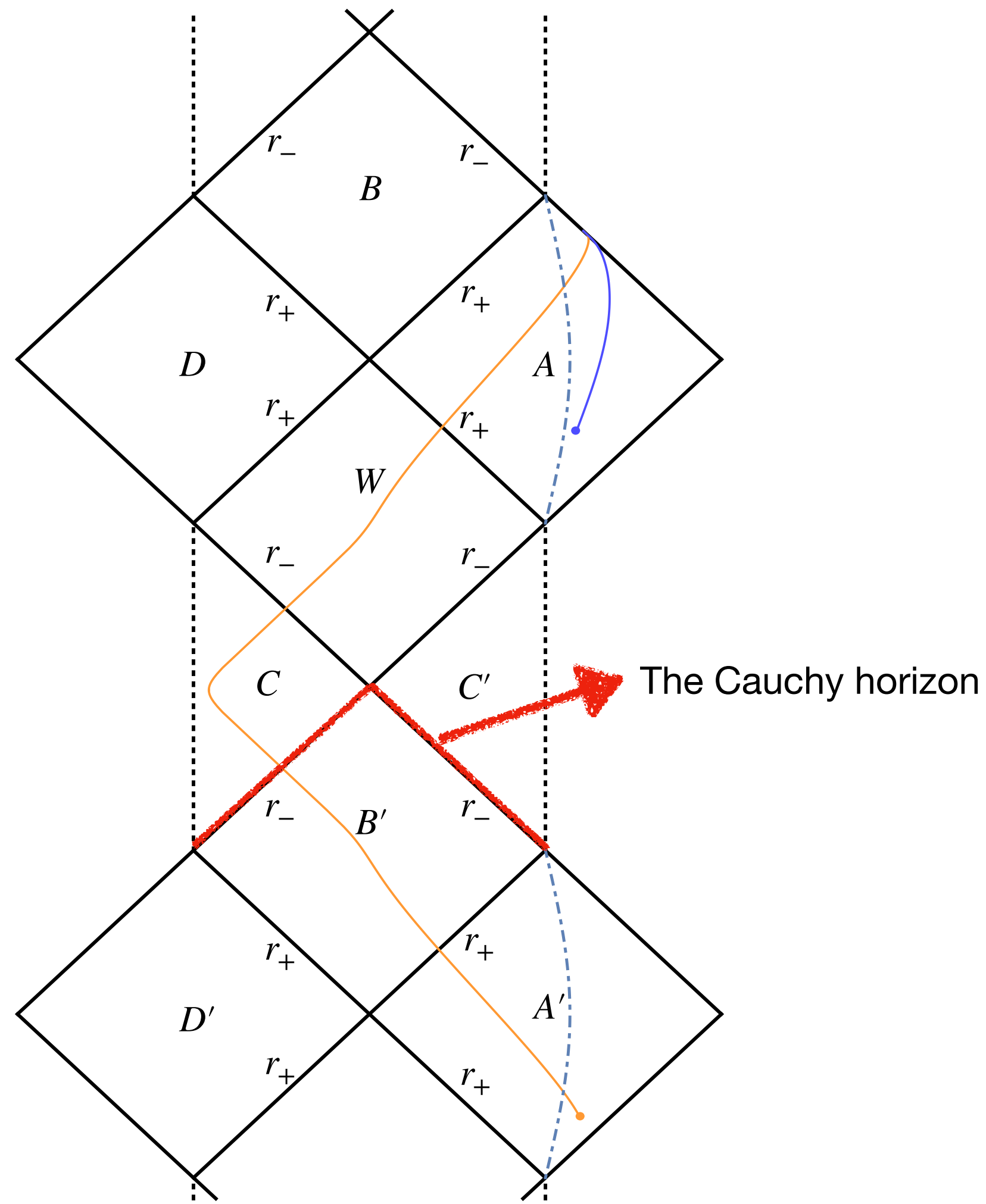
# Observing effect of LQBH

[CZ, Ma, Yang 23']



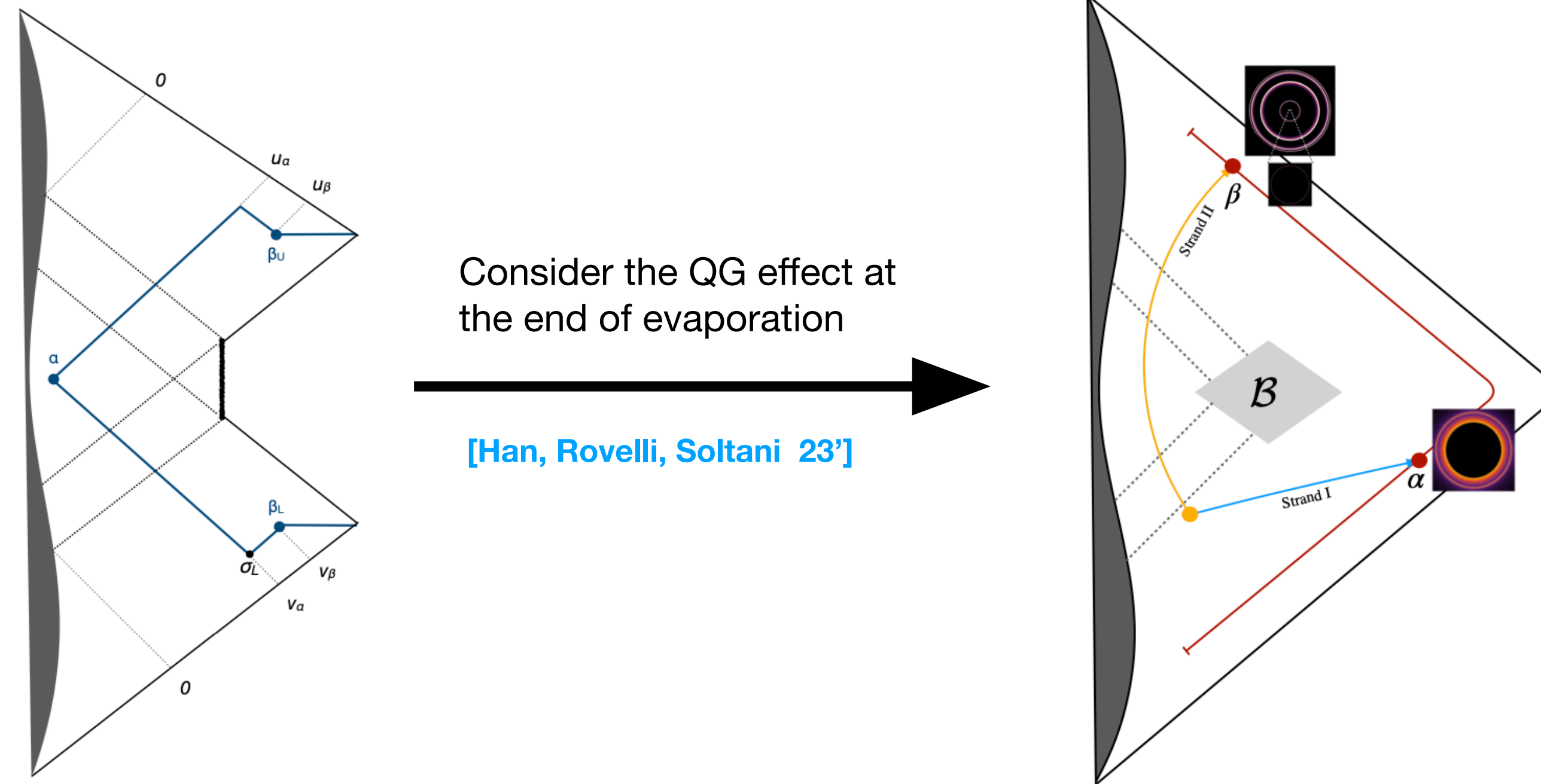
[see Cao, Li, Liu, Zhou 24' for similar work in regular BH]

# A new model



While the spacetime offers distinct advantages, it is not without debates: The existence of Cauchy horizon implies that the spacetime could be unstable under perturbation [\[Cao, Li et.al. 23' and 24', Shao, CZ, et.al. \(2023\)\]](#)

# A new model



**In a new work, we compute the SF transition amplitude in the B-Region, and study the tunneling of geometry in SF model.**

$$\det(e_+) = -\det(e_-), \text{ tunneling between opposite orientations } [\text{ArXiv: 2404.02796}]$$

# Summary

**We briefly review the loop quantum BH modes:**

- **The first type is constructed in virtue of the homogeneity of Schwarzschild interior;**
- **The second type considers the interior and exterior as a whole and does quantization;**
- **The third type considers matter collapsing.**

**In these models, we have some exciting results, like BH-WH transition, discrete mass spectrum, Nariai limit of BH evolution etc. However, our answer on loop quantum BH hasn't formed a unique picture.**

**We come up with an observable effects to test quantum BH.**

**Thank you for your attention !**